

Introduction to Quantum Theory

Dr. Russell Herman
Physics and Physical Oceanography

$E = hf$

It's relative

1927 Solvay Conference

My head is spinning

I'm uncertain

Complementarity

I have 2 Nobels

First row: I. Langmuir, M. Planck, M. Curie, H. A. Lorentz, A. Einstein, P. Langevin, C. E. Guye, C. T. R. Wilson, O. W. Richardson.
Second row: P. Debye, M. Knudsen, W. L. Bragg, H. A. Kramers, P. A. M. Dirac, A. H. Compton, L. V. de Broglie, M. Born, N. Bohr.
Third row: A. Piccard, E. Henriot, P. Ehrenfest, E. Herzen, T. de Donder, E. Schrödinger, E. Verschaffelt, W. Pauli, W. Heisenberg, R. H. Fowler, L. Brillouin.

Syllabus

• Website: <http://people.uncw.edu/hermanr/qm/>

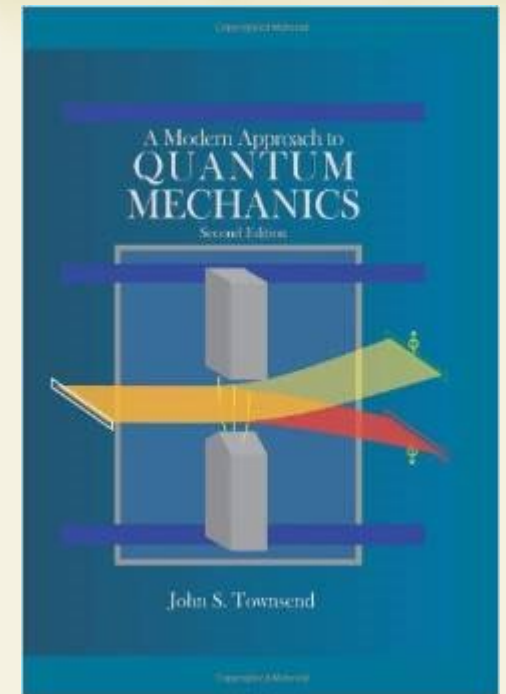
• Grades

Homework – 30%

Papers – 10%

3 Exams – 40%

Final – 20%



Required Text:

Townsend, *J. A Modern Approach to Quantum Mechanics*, 2nd Ed., 2012.

Other Readings:

Susskind, L. *Quantum Mechanics, The Theoretical Minimum*, 2014.

Feynman, R. C.,

The Feynman Lectures on Physics, Vol. III, 1965 and

QED: The Strange Theory of Light and Matter, 1988.

See also - <http://people.uncw.edu/hermanr/booklist.htm>

COVID-19

•Face to Face unless ...

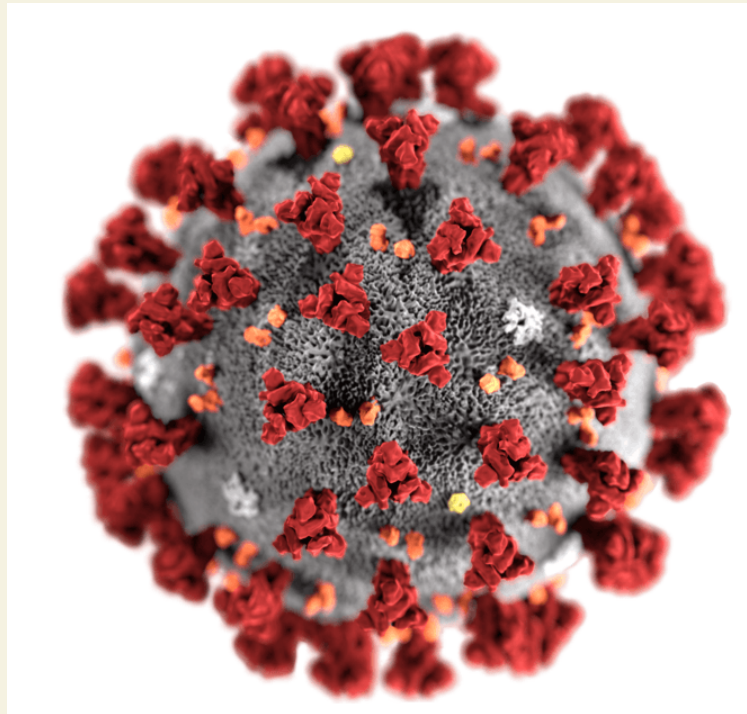
•<https://uncw.edu/coronavirus/>

–Social Distancing

–Face Coverings

–Wash Hands

•Office Hours



Following CDC Guidelines, UNC System directives, and out of mutual respect as outlined in the UNCW Seahawk Respect Compact, all faculty, staff, and students will wear face coverings while inside buildings. Students who are unprepared or unwilling to wear protective face coverings will not be permitted to participate in face-to-face sessions and will need to leave the building.

Noncompliant students will be referred to the Dean of Students for an Honor Code Violation. Any student who has a medical concern with wearing a face covering should contact the Disability Resource Center at (910) 962-7555.

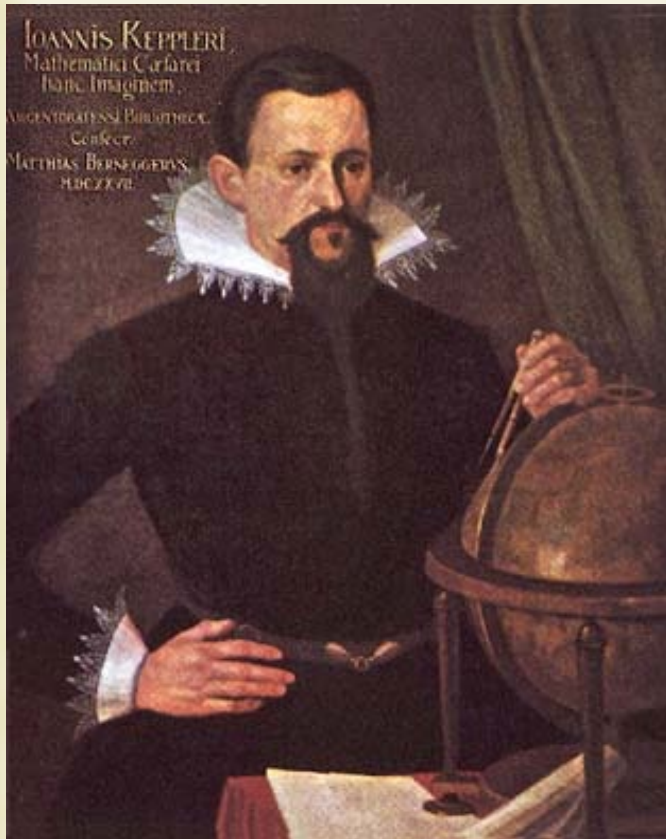
Students who experience COVID-19 symptoms should immediately contact the Abrons Student Health Center at (910) 962-3280.

The Rise of Classical Physics

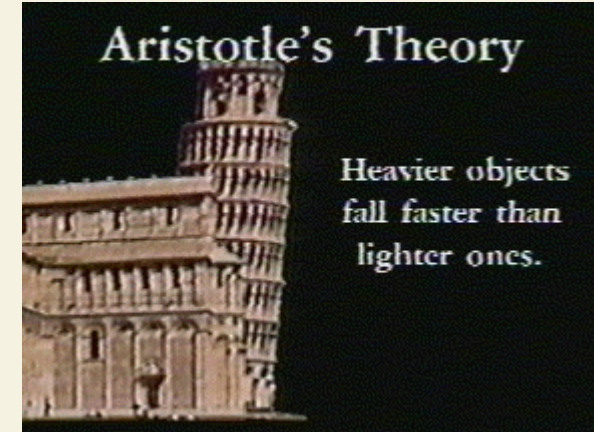
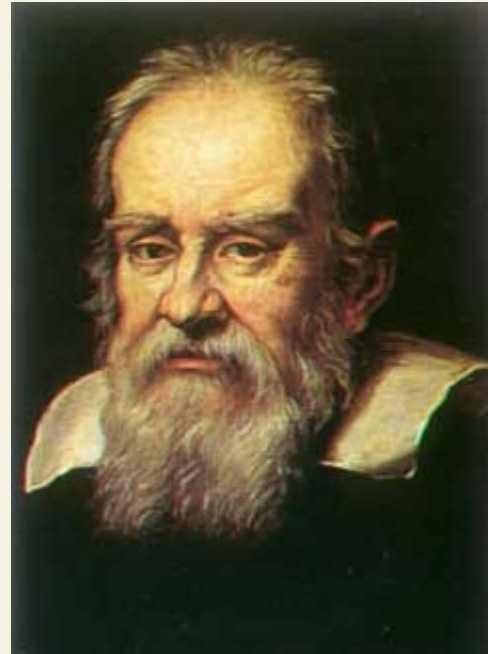
Scientific Deduction

Tycho Brahe
(1546-1601)

Johannes Kepler
(1571-1630)



Galileo Galilei
(1564-1642)



The Clockwork Universe

Sir Isaac Newton (1642-1727)

Principia (1687)

*Philosophiae Naturalis Principia
Mathematica (Mathematical
Principles of Natural Philosophy)*

Laws of Motion $dp/dt = \mathbf{F}$

Law of Gravitation

Kepler's Laws Explained

Calculus

... space is absolute ...

Determinism - Given \mathbf{F} , predict \mathbf{x} and \mathbf{v}



Unification

... the force responsible for bodies falling on the Earth is the same as that causing the moon to follow its orbit.

Reformulations of $F = ma$

Euler (1707-1783) **Variational Calculus**

D'Alembert (1717-1783) **Virtual Work**

Lagrange (1736-1813) **Lagrangian Mechanics**

Hamilton (1805-1865) **Hamiltonian Mechanics**

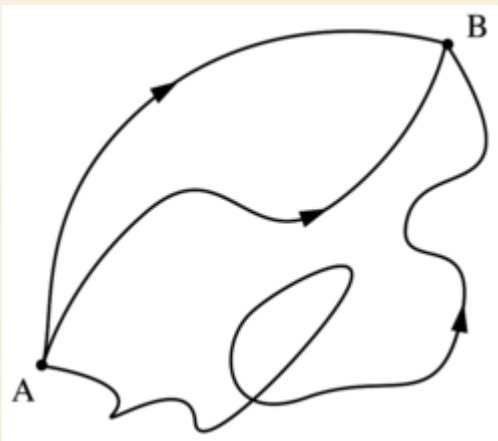
Define the action

$$S = \int_{t_1}^{t_2} L dt, \text{ for } L = T - V.$$

Require: $\delta S = 0$.

Then,

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) - \frac{\partial L}{\partial x} = 0.$$



Principles

Fermat's

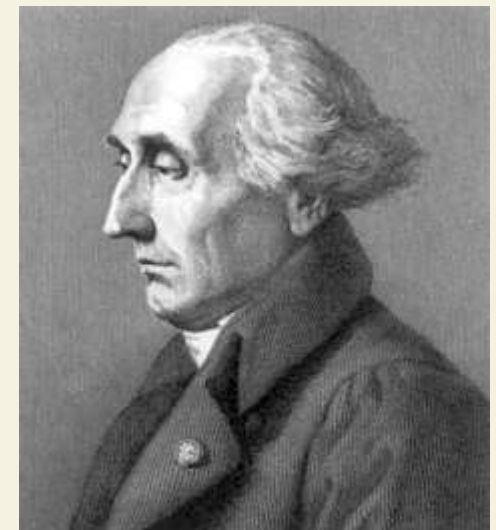
least time

d'Alembert's

virtual work

Hamilton's

least action



Hamilton's Formulation

Phase Space ($q=x$, $p= m dx/dt$),

Initial (q, p) + 2nd Law \Rightarrow Motion for all t

Ex: Free particle, $p = \text{const}$

Harmonic Oscillator,

Energy Conservation $E=\text{const}$

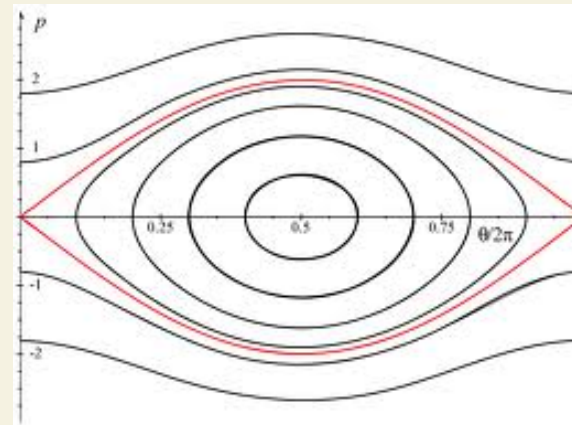
$$E = \frac{p^2}{2m} + \frac{1}{2}kq^2$$



Hamiltonian $H(p,q) = T(p) + V(q)$

Hamilton's Eqns \Leftrightarrow Newton's Laws

$$\begin{aligned} \frac{dq}{dt} &= \frac{\partial H(q, p)}{\partial p} \\ \frac{dp}{dt} &= -\frac{\partial H(q, p)}{\partial q} \end{aligned}$$



Electricity and Magnetism

Magnetism

Lode stones

Compasses

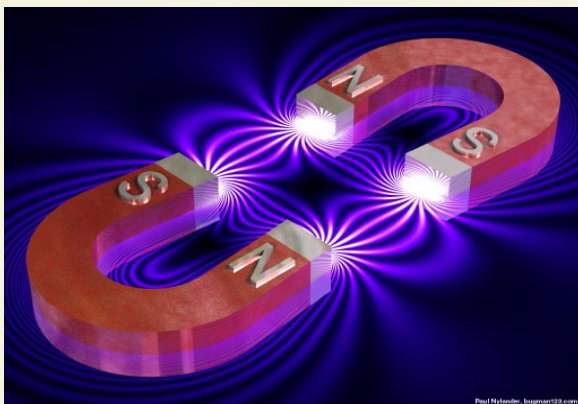
William Gilbert (1544-1603)

Thomas Browne (1605-1682)

Benjamin Franklin (1706-1790)

Luigi Galvani (1737-1798)

Alessandro Volta (1745-1827)



Electromagnetism

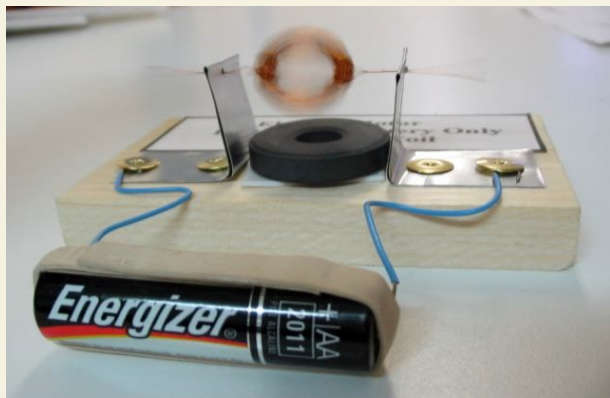
André-Marie Ampère, (1775 - 1836)

Hans Oersted (1777-1851)
current deflects compass needles

Georg Simon Ohm (1789-1854)

Joseph Henry (1797-1878)
electromagnetic induction, first motor,
telegraph

Michael Faraday (1791-1867)
electrolysis, motors, induction coils, ...



Electromagnetic Waves

James Clerk Maxwell (1831-1879)

- Theory of electromagnetism.
- Predicted the electromagnetic waves.
- Electromagnetic waves travel at

$$c = 299,792,458 \text{ m/s} = 186,000 \text{ mi/s}$$

Heinrich Hertz (1857-1894)

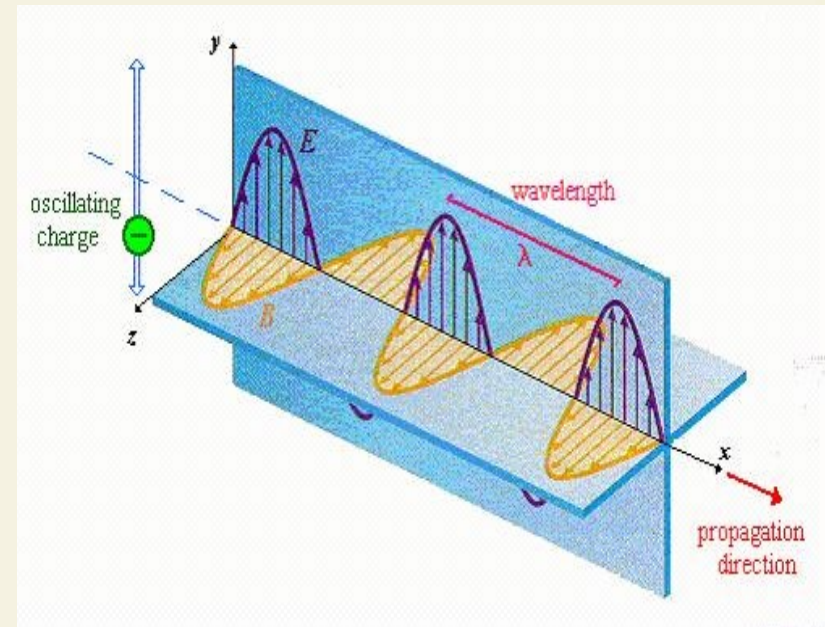
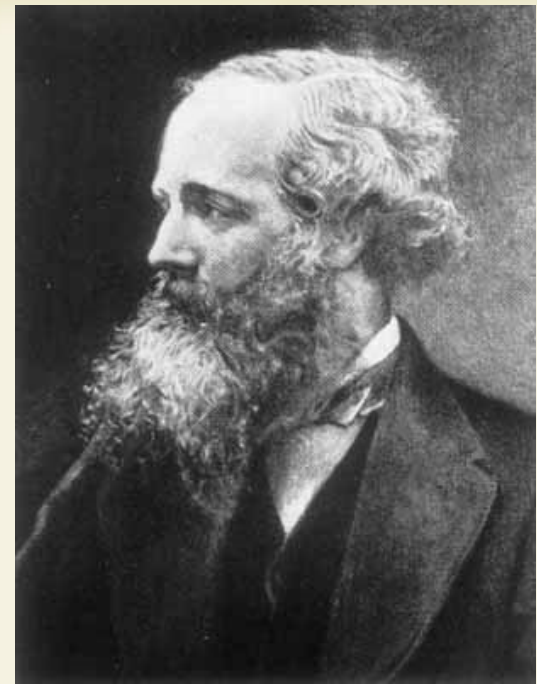
- sent the first radio waves (1888)

What is the medium?

Luminiferous Aether

Michelson-Morley (1887)

- could not detect it.



Maxwell's Equations

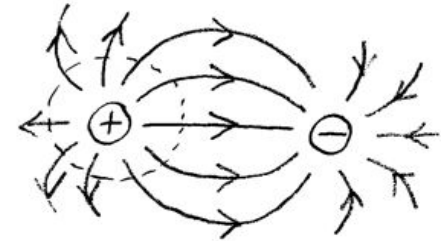
Not SI units!

JAMES CLERK MAXWELL'S EQUATIONS

$$\nabla \cdot \vec{E} = 4\pi\rho$$

DIVERGENCE OF \vec{E} CHARGE DENSITY

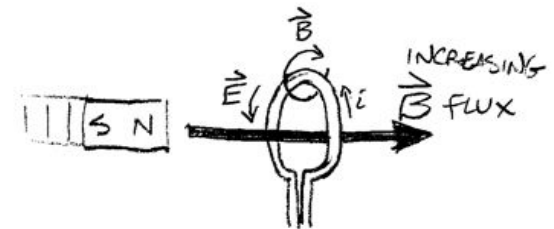
\vec{E} DIVERGES OUT FROM POSITIVE CHARGES AND IN TOWARD NEGATIVE CHARGES. THE TOTAL FLUX OF \vec{E} THROUGH ANY CLOSED SURFACE IS PROPORTIONAL TO THE CHARGE INSIDE.



$$\nabla \times \vec{E} = -\frac{1}{c} \frac{d\vec{B}}{dt}$$

CURL OF \vec{E} SPEED OF LIGHT RATE \vec{B} IS CHANGING

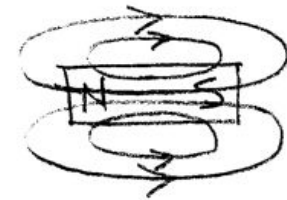
\vec{E} CURLS AROUND CHANGING \vec{B} FIELDS (FARADAY'S LAW) IN A DIRECTION THAT WOULD MAKE A CURRENT THAT WOULD PRODUCE A \vec{B} FIELD TO OPPOSE THE CHANGE IN \vec{B} FLUX (LENZ'S LAW).



$$\nabla \cdot \vec{B} = 0$$

DIVERGENCE OF \vec{B}

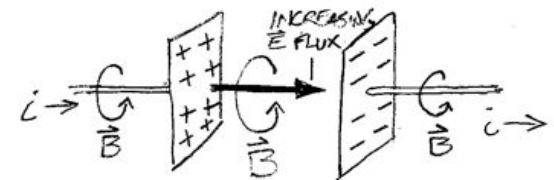
\vec{B} NEVER DIVERGES. IT JUST LOOPS AROUND ON ITSELF.



$$\nabla \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{d\vec{E}}{dt}$$

CURL OF \vec{B} SPEED OF LIGHT CURRENT DENSITY RATE \vec{E} IS CHANGING

\vec{B} CURLS AROUND CURRENTS AND CHANGES IN \vec{E} FIELDS



Gaussian Units

Unit Conversions



		Conversion	SI
Distance	cm	10^{-2}	m
Mass	g	10^{-3}	kg
Time	s	1	s
Force	dyne	10^{-5}	N
Energy	erg	10^{-7}	J
Power	erg/s	10^{-7}	W
Charge	esu	3.336×10^{-10}	C
Electric Potential	statvolt	299.79	V
Magnetic Field	Gauss	10^{-4}	T

- $1 \text{ eV} = 1.6022 \times 10^{-12} \text{ erg} = 1.602 \times 10^{-19} \text{ J}$
- $1 \text{ Ry} = 13.6057 \text{ eV}$ (ionization energy of hydrogen)
- $1 \text{ C} = 2.9979 \times 10^9 \text{ esu}$, $1 \text{ statcoul} = 1 \text{ esu}$
- $1 \text{ \AA} = 10^{-10} \text{ m}$
- $1 \text{ eV}/c^2 = 1.7827 \times 10^{-36} \text{ kg}$
- $(\mu_0 \epsilon_0)^{-1/2} = 299,792,458 \text{ m/s}$; $(\mu_0/\epsilon_0)^{1/2} \approx 377 \Omega$

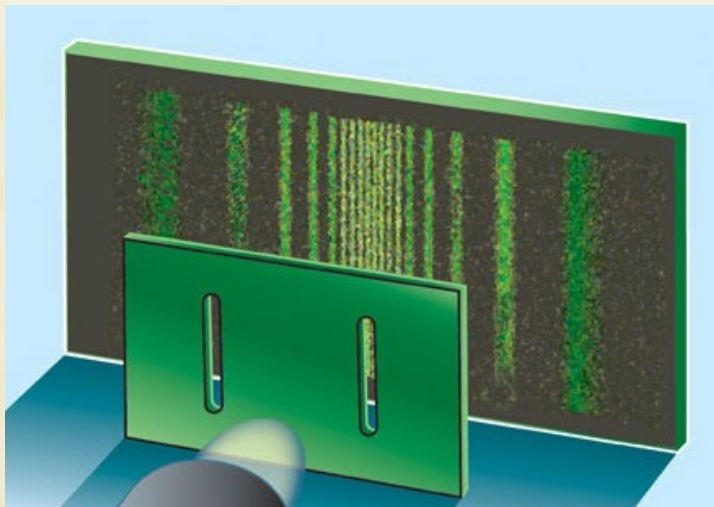
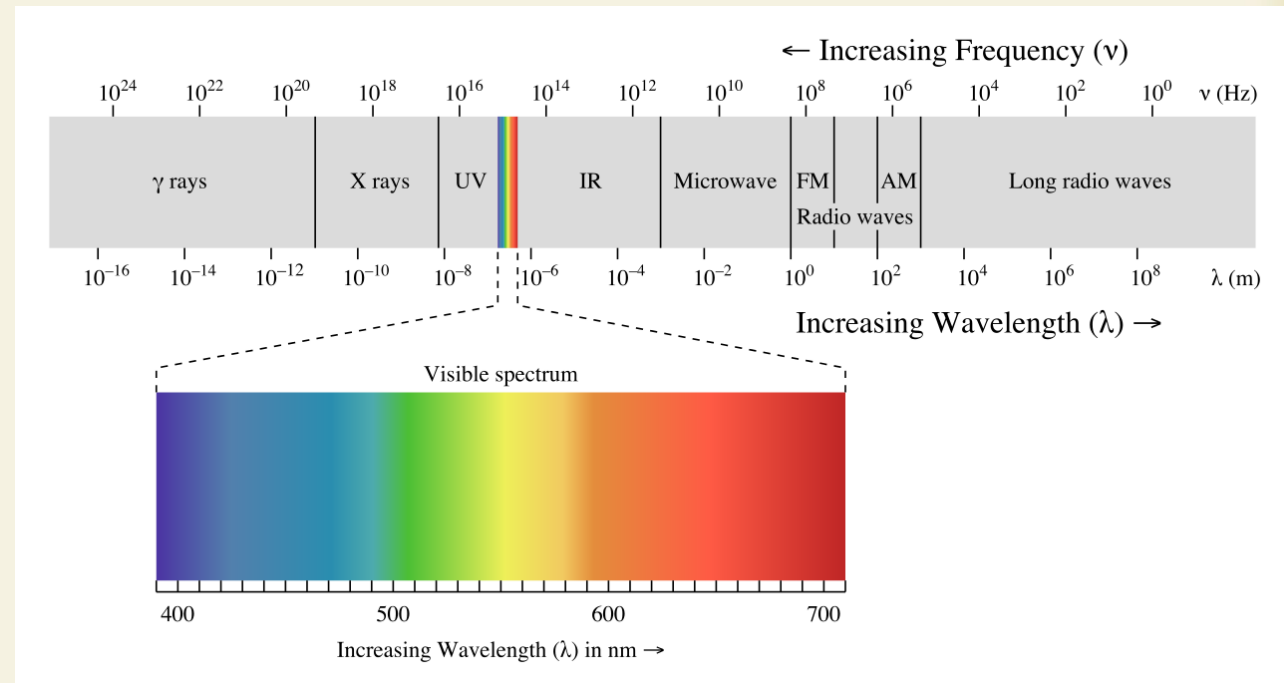
WAVES

What are waves?

- **Characteristics**
 - Wavelength, Frequency, Wavespeed
- **Behavior**
 - Superposition, Interference. Diffraction



$$\frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = \nabla^2 E$$



Spectroscopy

Ionized gas gives off radiation

Johann Balmer 1885

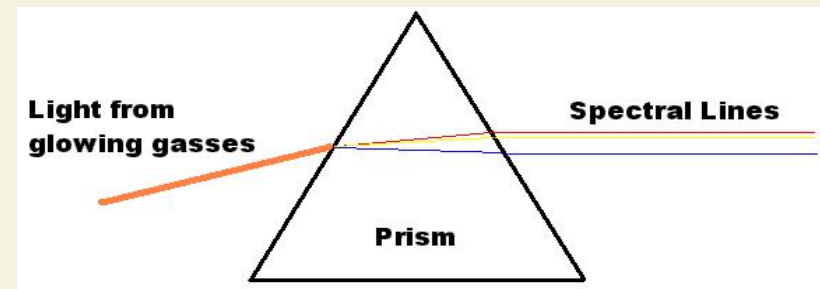
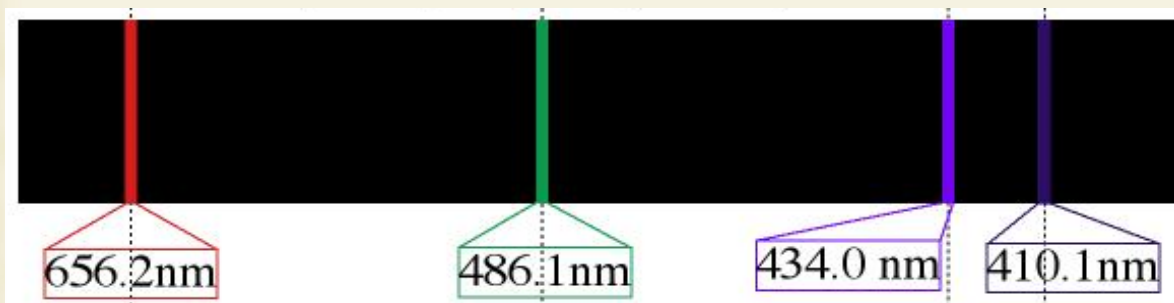
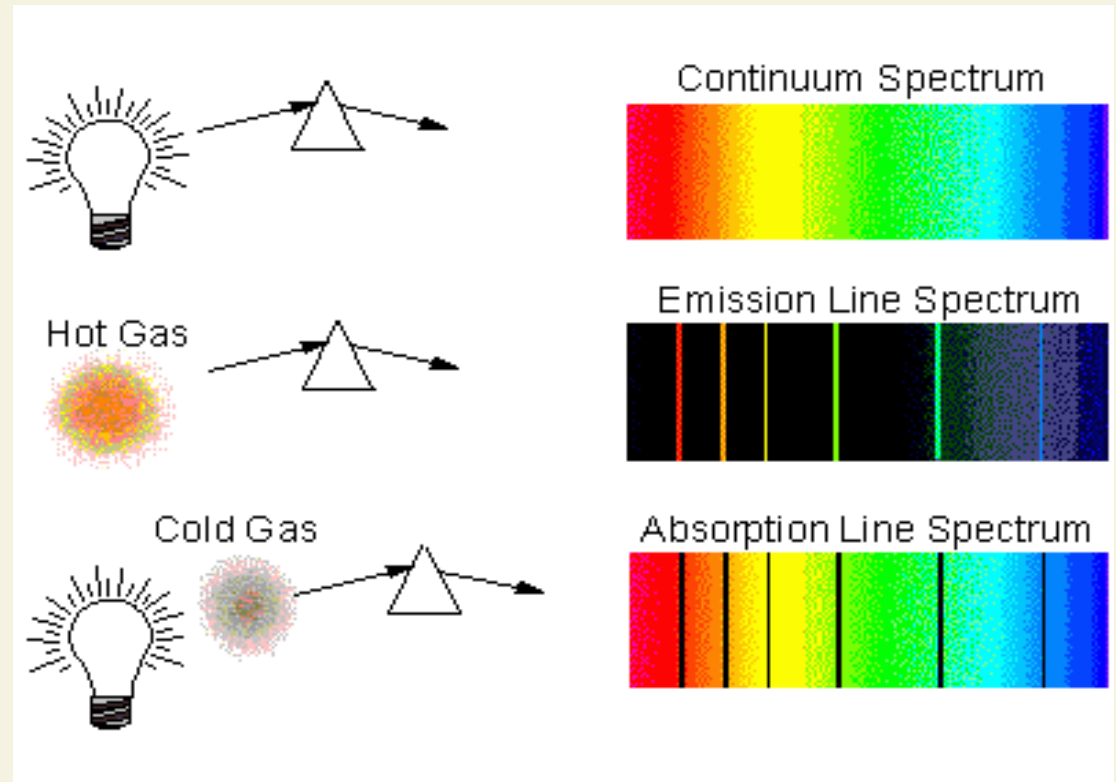
Spectral Lines: Hydrogen
410, 434, 486, 656 nm

Empirical Formula:

$$\lambda = R \left(\frac{1}{4} - \frac{1}{n^2} \right)$$

Predicted 5th-7th lines

Lyman and Paschen Series



Laws of Thermodynamics

Engines: Watt, Carnot, Kelvin, Clausius, Carnot

James Joule (1818-1889)

Mechanical Equivalent of Heat

Laws of Thermodynamics

1. Adding heat energy or doing work on a body increases internal energy.
(Energy conservation)
2. A body will not spontaneously get hotter.
(Entropy and the Arrow of Time)

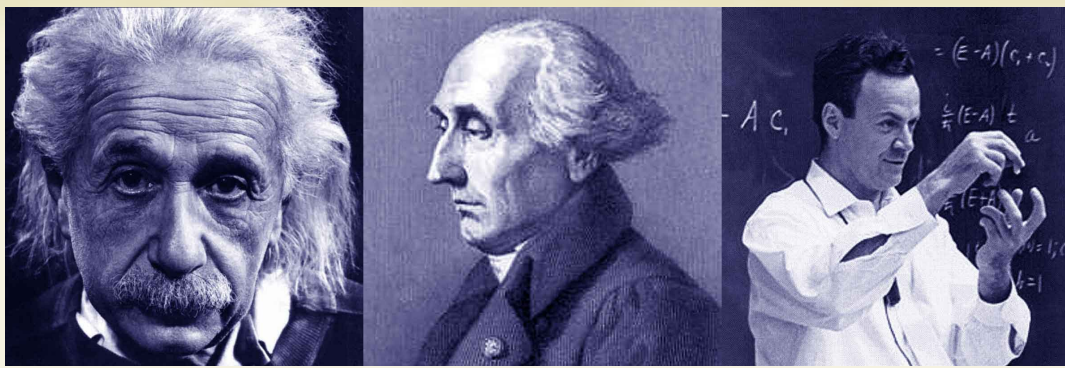
Joseph Stefan (1835-1893) and Ludwig Boltzmann (1844-1906)

Heated bodies Radiate - Stefan-Boltzmann Law

Radiation from blackbody proportional to T^4 .

$$P = e\sigma AT^4$$

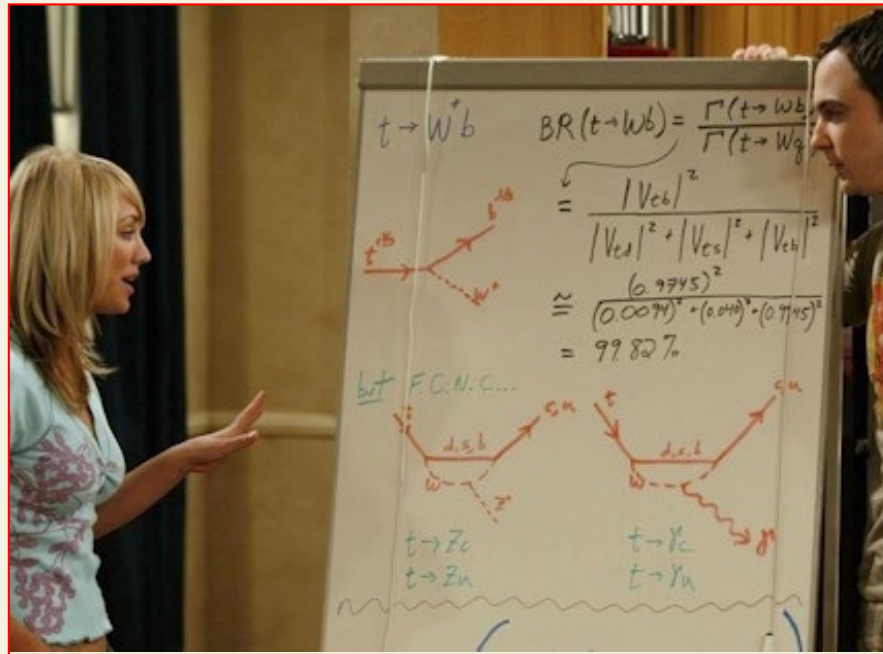
Maxwell-Boltzmann Statistical Mechanics – *Bah Humbug!*



$$H|\psi\rangle = i\hbar \frac{\partial}{\partial t} |\psi\rangle$$

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

This is not your grandfather's physics

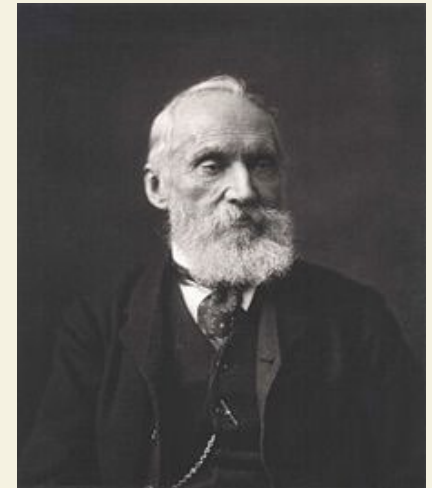


$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

$$\left[i\hbar \gamma^\mu \partial_\mu - mc \right] \psi = 0$$

Physics Revolutions

William Thomson, (1824 – 1907)
1st Baron Kelvin (**Lord Kelvin**)



“There is nothing new to be discovered in physics now. All that remains is more and more precise measurement.” - 1900

- 1895 **Wilhelm Röntgen** discovers **X-rays**. NP 1901
- 1896 **Henri Becquerel** discovers **radioactivity** NP 1903
- 1898 **Marie and Pierre Curie** separate radioactive elements NP 1903
- 1897 **Joseph (J.J.) Thomson** measures electron, NP 1906



“plum-pudding” model of atom”



Radioactivity and the Atom

1897 - J.J. Thomson discovers the **electron** NP 1906.

1898 - Marie and Pierre Curie discover the first radioactive elements:
radium and polonium NP 1903.

1899 - Ernest Rutherford divided radiation into **alpha and beta rays** NP 1908.

1900 - Pierre Curie observes **gamma rays**.

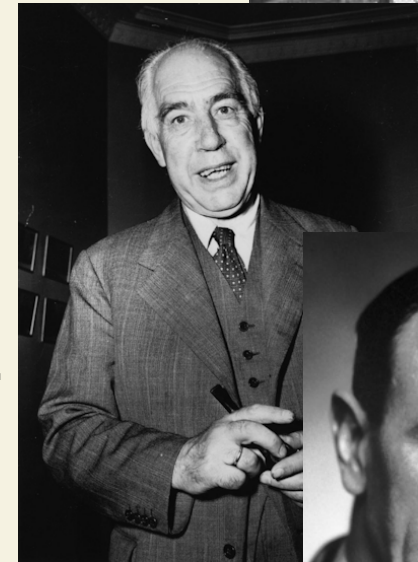
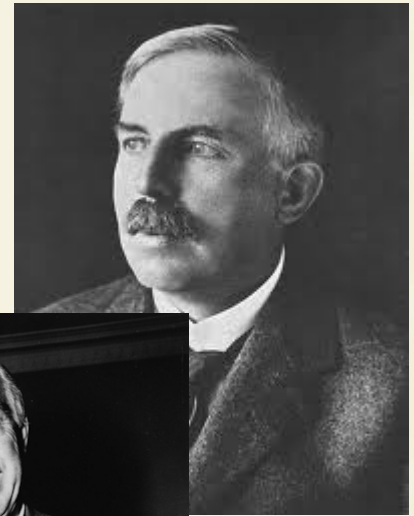
1911 - Ernest Rutherford discovers the **atomic nucleus**

1913 - Niels Bohr introduces the **first atomic model**, NP 1922
the mini solar system.

1913 - Hans Geiger invents **counter** for measuring radioactivity.

1920 - Ernest Rutherford discovered and named the **proton**.

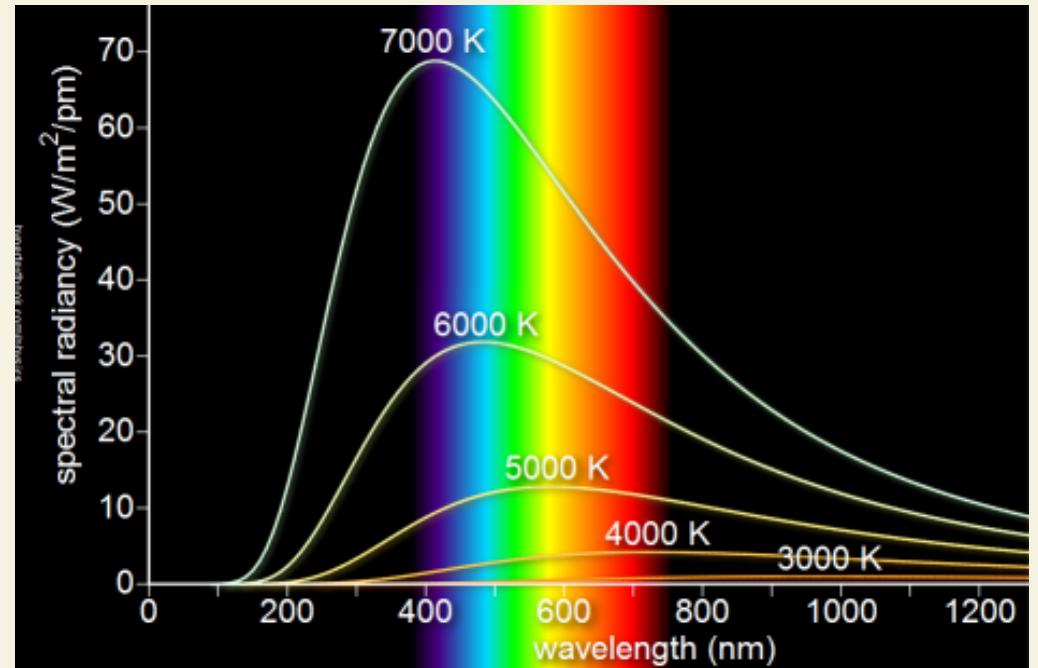
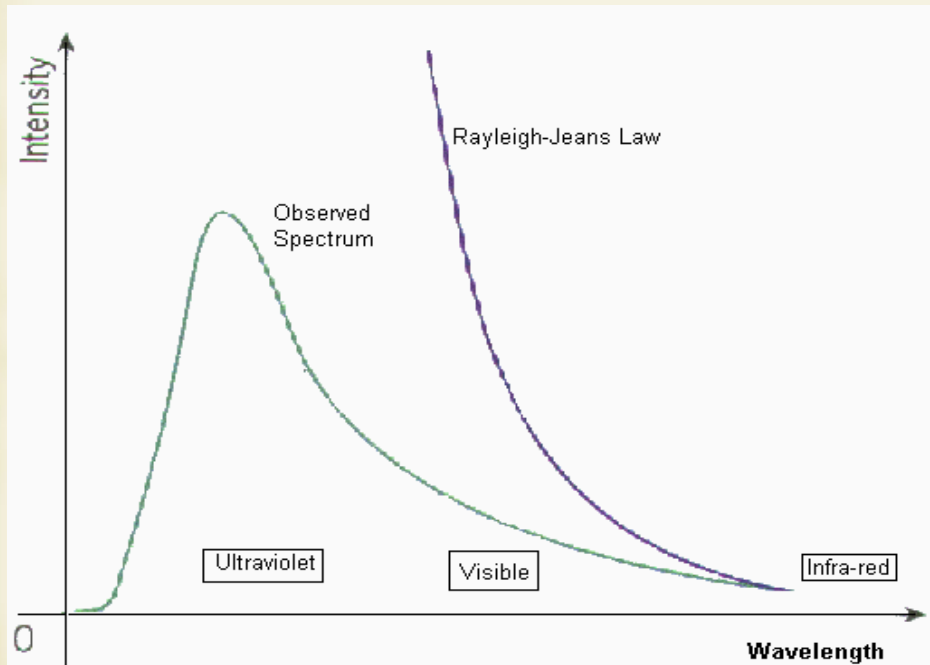
1932 - James Chadwick discovers the **neutron**. NP 1935



Blackbody Spectrum

Blackbody - a theoretical object that absorbs 100% of the radiation that hits it.

Wien's Law (1896) ^{NP 1911} and **Rayleigh** ^{NP 1904} - **Jeans Law** (1900)



Ultraviolet Catastrophe “... when you turn on your toaster, you are instantly fried by a massive gamma ray burst, since your little blackbody toaster should emit infinite energy at the shortest wavelengths.”

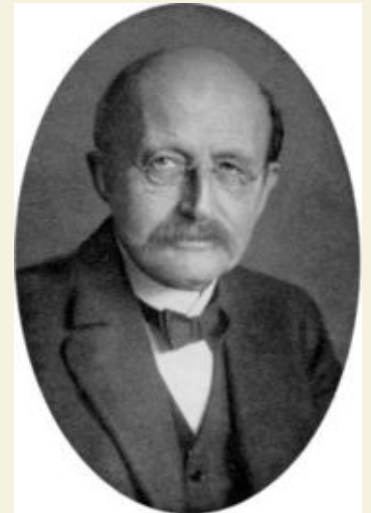
Quantum Theory

Max Planck NP 1918

(Karl Ernst Ludwig Max Planck 1858-1947)

oscillators can only vibrate at discrete frequencies:

$$E_n = n(hf), n = 1, 2, 3 \dots$$

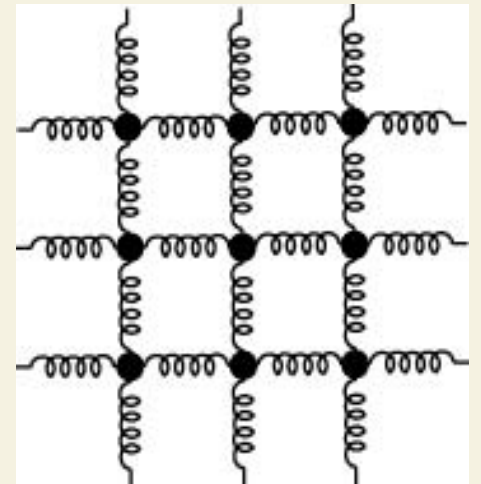


Thus, the energy difference

$$\Delta E = hf,$$

where Planck's constant is given by

$$h = 6.63 \times 10^{-34} \text{ Js}$$

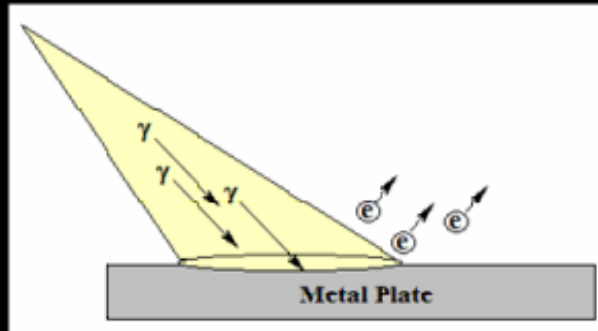


Albert Einstein - 1905

Photoelectric Effect

Light can cause currents

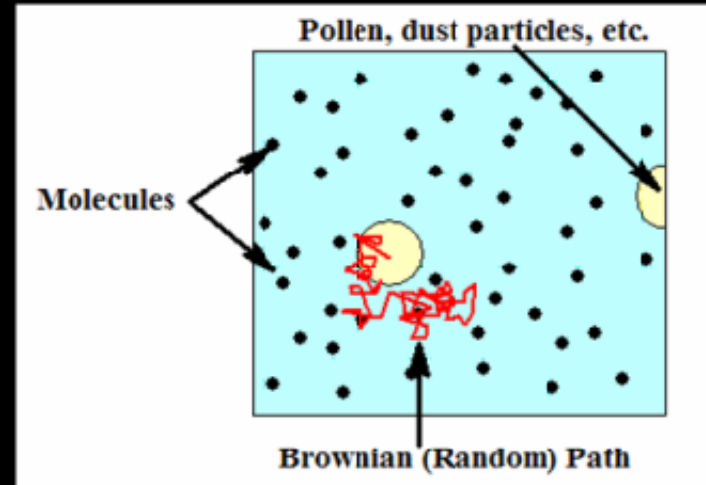
- Electrons can be ejected from irradiated metal plates.
- Light can be act like either particles (quanta) or waves.
- Extended Planck's ideas of energy quantization.
- Lead to explanation of electromagnetic spectra,
- Lead to the development of lasers, transistors and other applications.



Brownian Motion

the random movement of particles suspended in a fluid

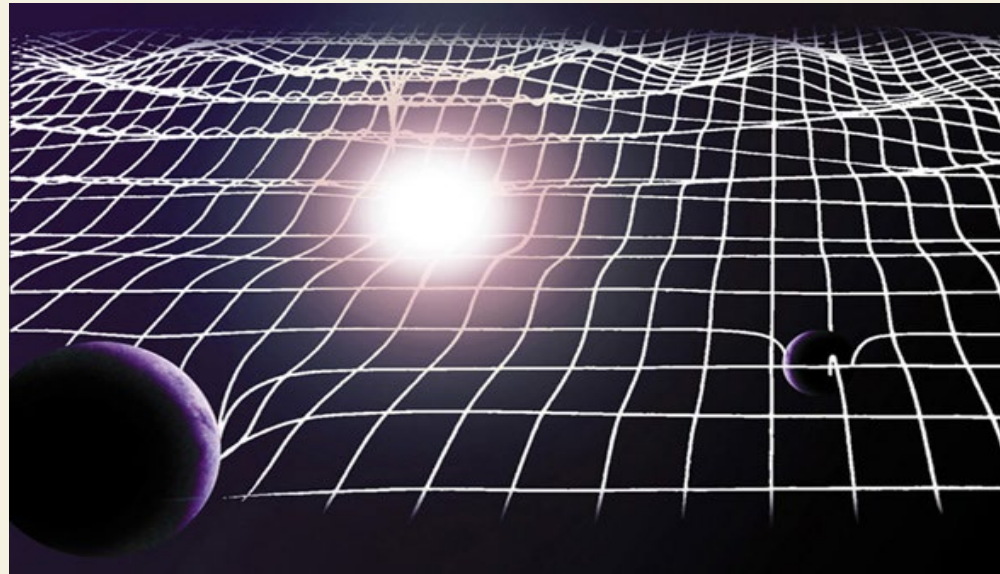
- Explained the observations credited to Robert Brown, 1827
- Predicted molecular motion and size through the effects of collisions with larger particles
- Einstein's work lead to an acceptance of molecular theory



Paradigm Shifts

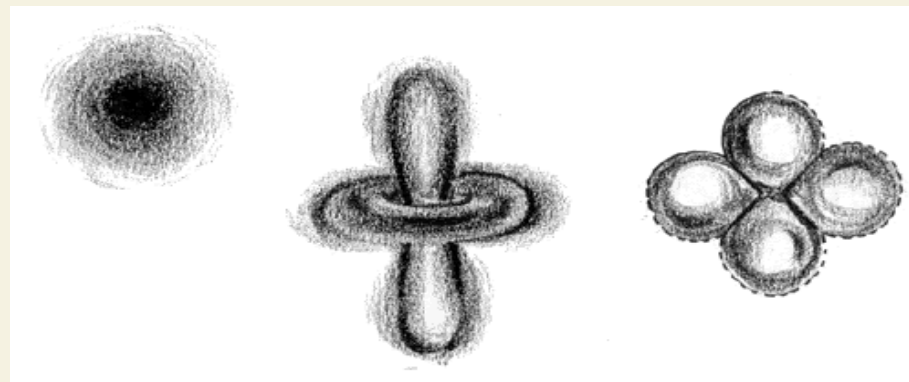
Relativity

Space and Time not absolute, not Euclidean



Quantum Mechanics

Loss of Determinism



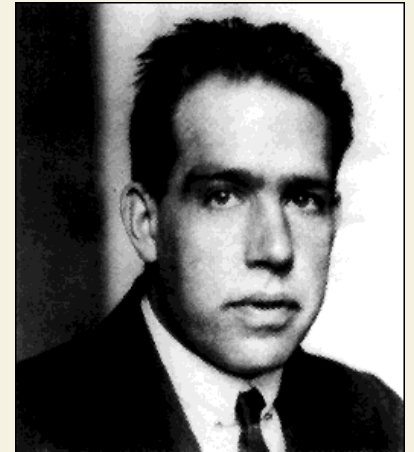
Bohr's Atom - 1913

Niels Bohr (1885-1962)

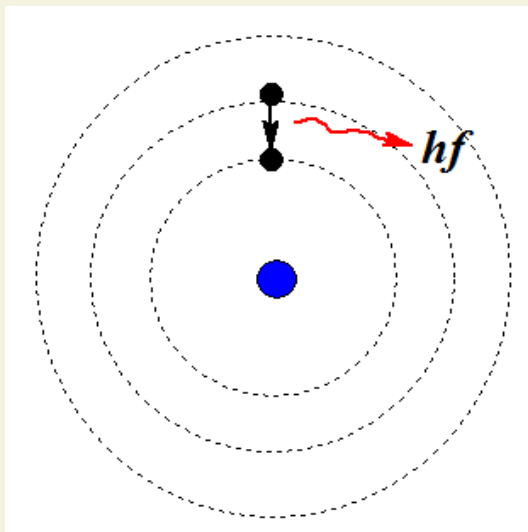
Accelerating electrons radiate at specific energies.

Are atoms stable?

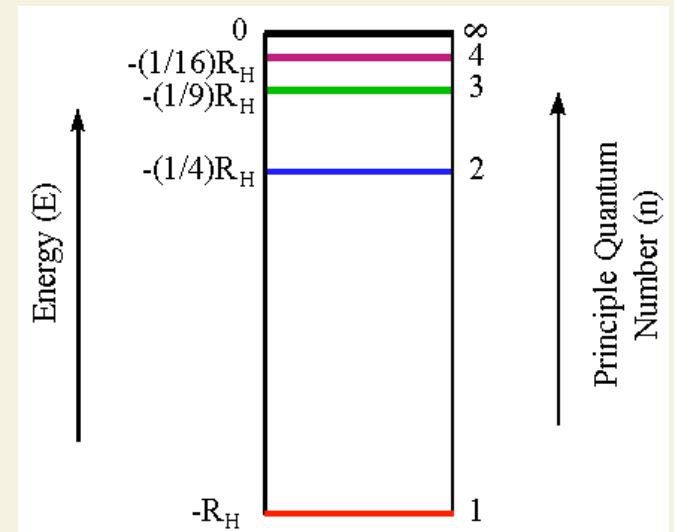
Assume angular momentum is quantized.



Niels Bohr

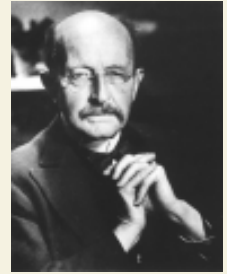


$$E_n = R_H \left(\frac{1}{m^2} - \frac{1}{n^2} \right)$$

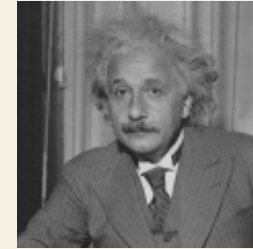


Early History - Quantum Mechanics

1900 - Planck Explains Blackbody Radiation NP 1906



1905 - Einstein - the Photoelectric Effect, Photons NP 1921



1913 - The Bohr Model for Hydrogen NP 1922



1916 - Confirmation of photon, Millikan NP 1923

1922 – Stern-Gerlach Experiment NP 1943

1923 - Compton NP 1927 Effect – X-Ray Scattering



1924 – de Broglie NP 1929 - Particles Behave Like Waves



1925 - Matrix Mechanics – Heisenberg NP 1932, Born NP 1954, Jordan
- Pauli Principle NP 1935



1926 - Wave Mechanics - Schrödinger NP 1933



1927 - The Uncertainty Principle – Heisenberg
- Davisson NP 1937 –Germer, Thomson NP 1937 - Verified de Broglie's idea

1928 - Relativistic Quantum Mechanics - Dirac NP 1933

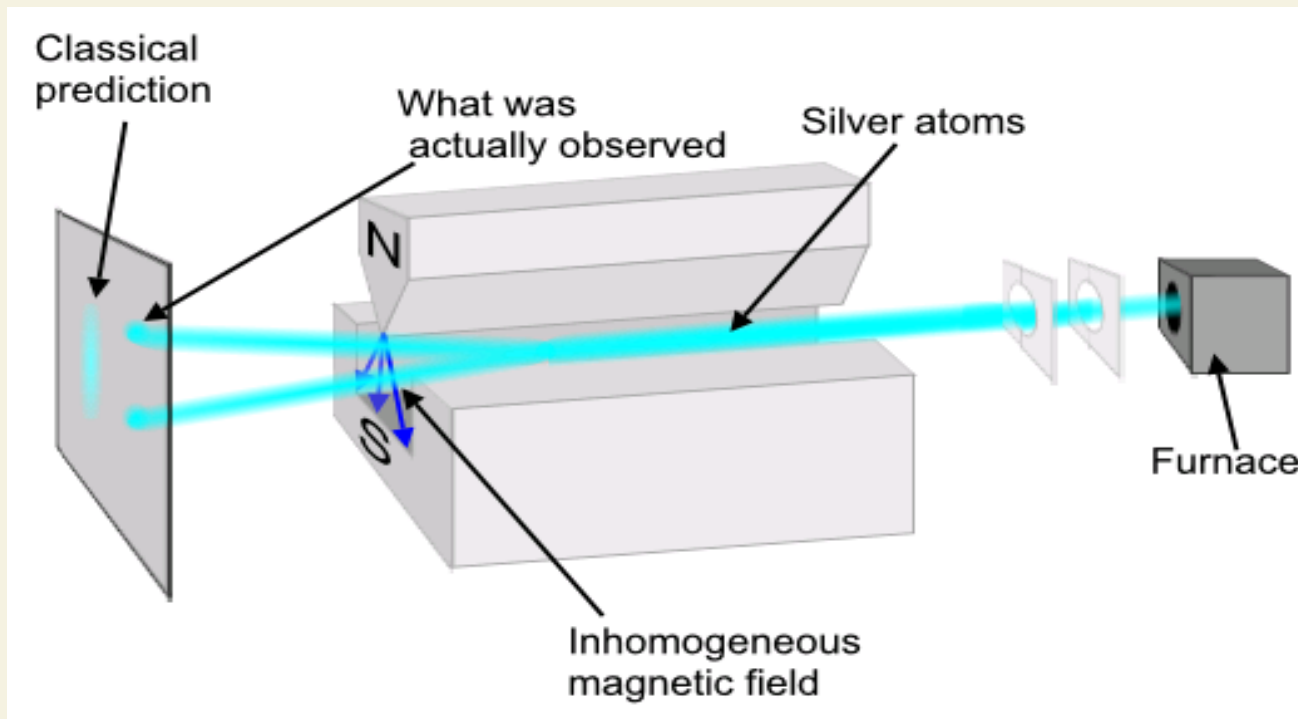
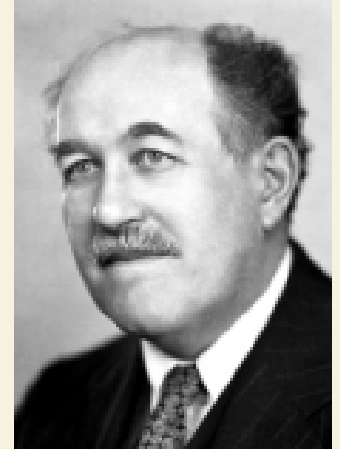


Stern-Gerlach Experiment - 1922

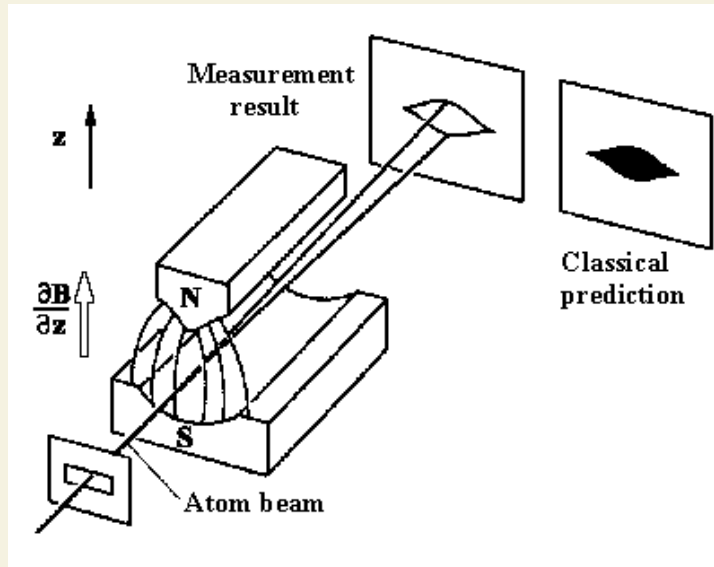
Otto Stern (1888-1969)

Walther Gerlach (1889-1979) – Was it the cigar?

Ag atoms sent through inhomogeneous magnetic field.



Stern-Gerlach Results



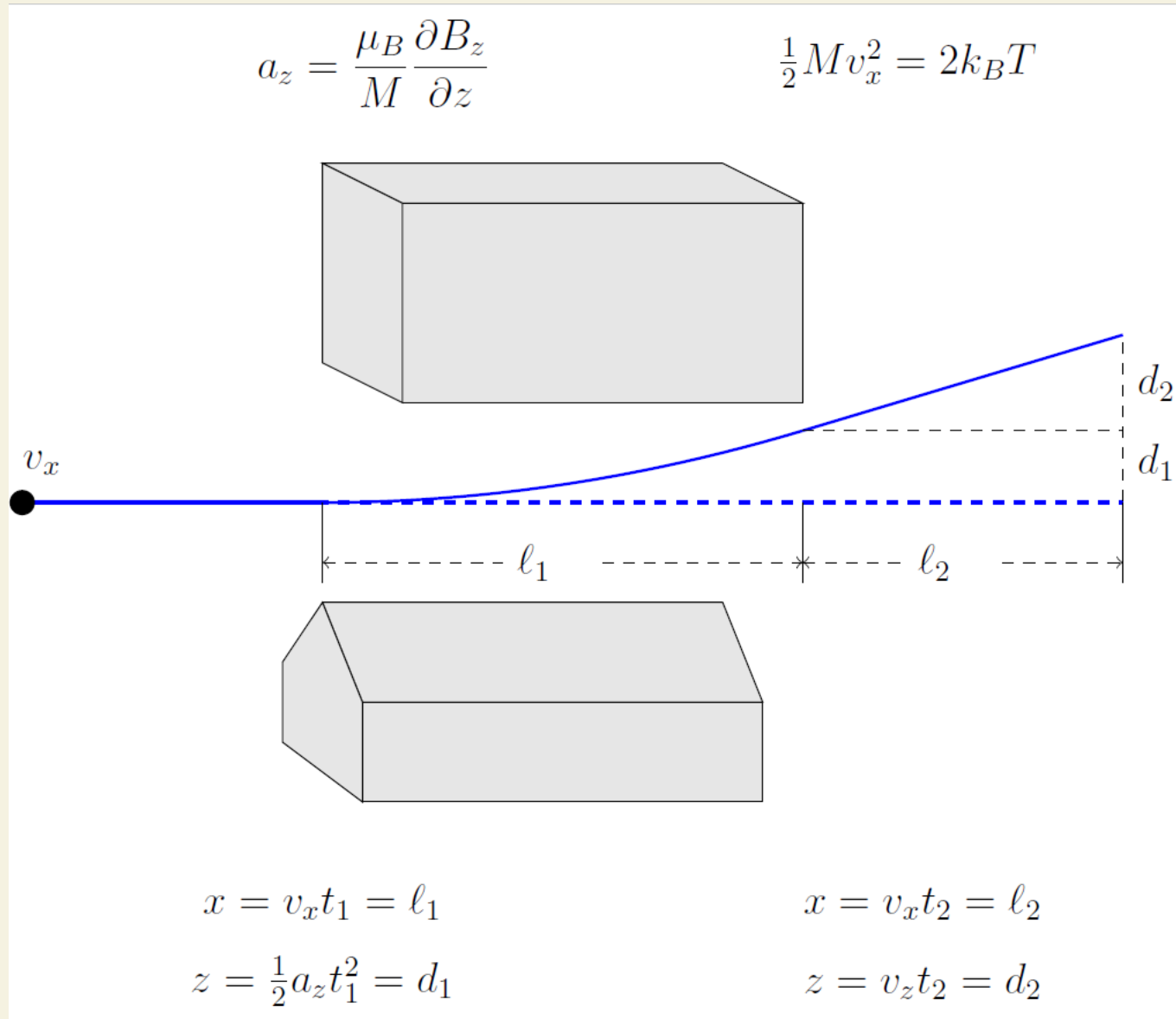
Test of classical vs quantum theory of angular momentum (L).

$L = 0$ or $L = 1$ – No splitting for $L = 0$?

Uhlenbeck and Goldsmit (1925,1926) proposed **intrinsic spin**.

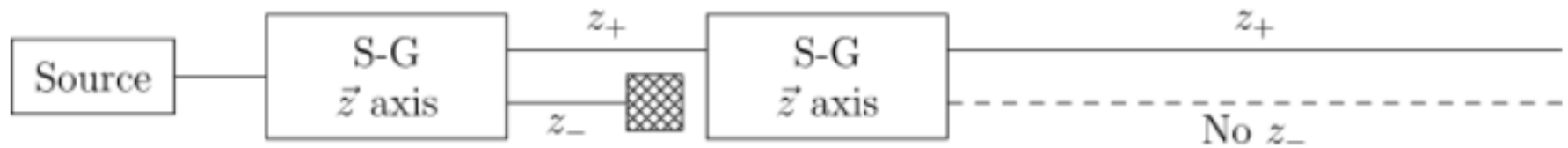
- to explain the anomalous Zeeman effect,
(the splitting of spectral lines in a magnetic field).

Stern-Gerlach Particle Path – Problem 1

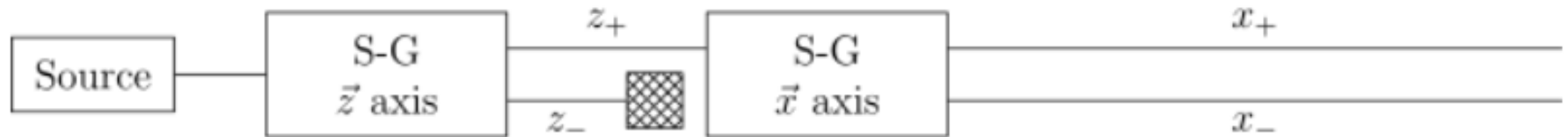


Stern-Gerlach Thought Experiments

1



2



3

