

# Introduction to Quantum Theory

*Dr. Russell Herman*  
*Physics and Physical Oceanography*

$E = hf$

It's relative

1927 Solvay Conference

My head is spinning

I'm uncertain

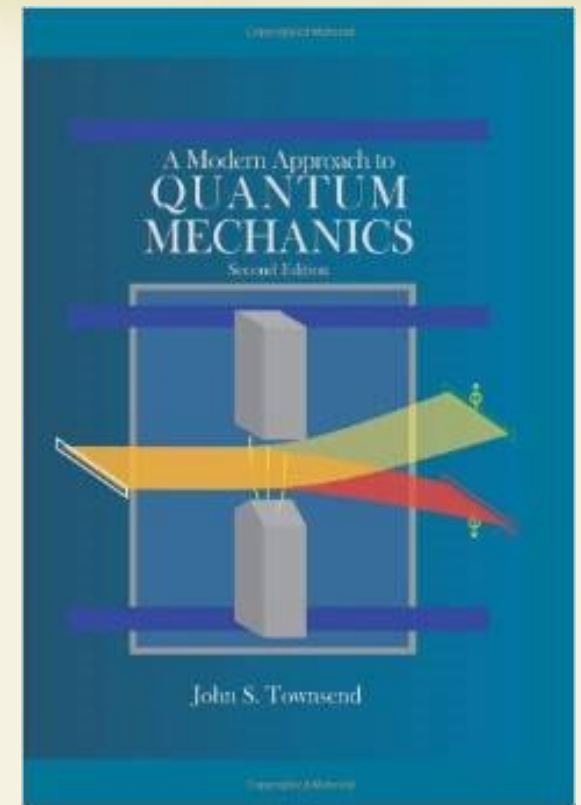
Complementarity

I have 2 Nobels

First row: I. Langmuir, M. Planck, M. Curie, H. A. Lorentz, A. Einstein, P. Langevin, C. E. Guye, C. T. R. Wilson, O. W. Richardson.  
Second row: P. Debye, M. Knudsen, W. L. Bragg, H. A. Kramers, P. A. M. Dirac, A. H. Compton, L. V. de Broglie, M. Born, N. Bohr.  
Third row: A. Piccard, E. Henriot, P. Ehrenfest, E. Herzen, T. de Donder, E. Schrödinger, E. Verschaffelt, W. Pauli, W. Heisenberg, R. H. Fowler, L. Brillouin.

# Syllabus

- Website: <http://people.uncw.edu/hermanr/qm/>
- Grades
  - Homework – 30%
  - Papers – 10%
  - 3 Exams – 40%
  - Final – 20%
- Office Hours: MTWRF, 10-11 AM  
Sartarelli Hall 2007J



## Required Text:

Townsend, J. *A Modern Approach to Quantum Mechanics*, 2nd Ed., 2012.

## Other Readings:

Susskind, L. *Quantum Mechanics, The Theoretical Minimum*, 2014.

Feynman, R. C.,

*The Feynman Lectures on Physics*, Vol. III, 1965 and

*QED: The Strange Theory of Light and Matter*, 1988.

See also - <http://people.uncw.edu/hermanr/booklist.htm>

# Time for Some Background

## The Rise of Classical Physics

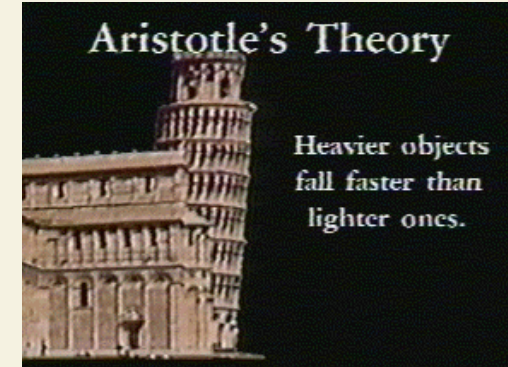
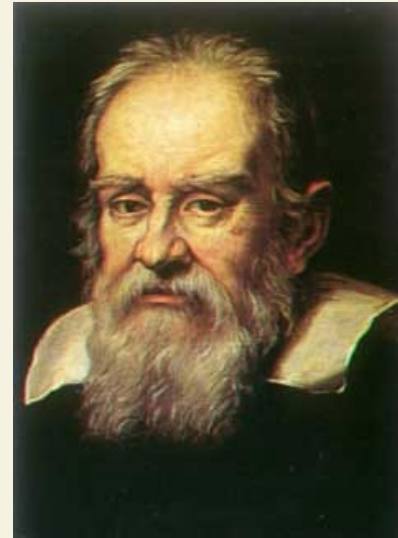


# The Emergence of Physics - 1609

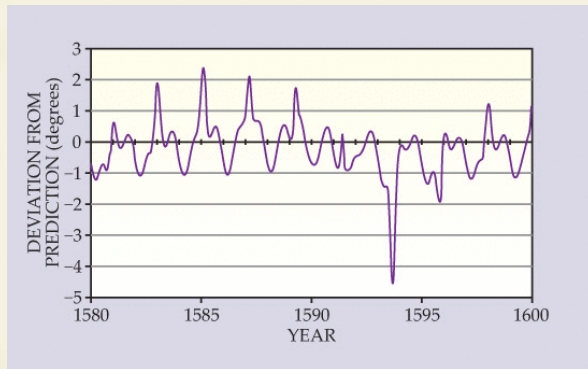
**Nicolaus Copernicus (1473-1543)**



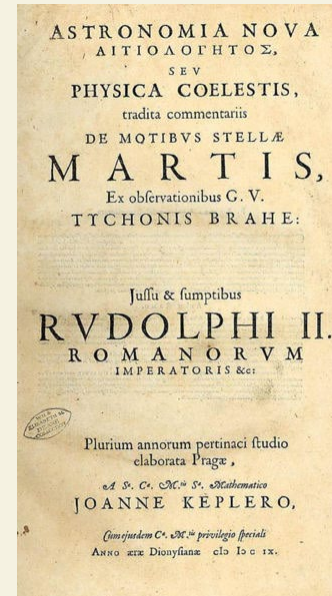
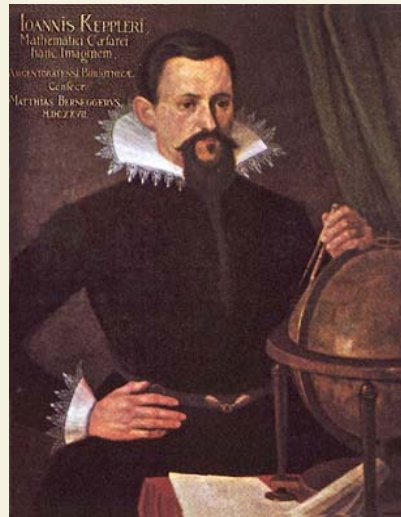
**Galileo Galilei (1564-1642)**



**Tycho Brahe (1546-1601)**



**Johannes Kepler (1571-1630)**



**The great Martian catastrophe and how Kepler fixed it**





# The Clockwork Universe

**Sir Isaac Newton (1642-1727)**

**Principia (1687)**

*Philosophiæ Naturalis Principia  
Mathematica (Mathematical  
Principles of Natural Philosophy)*

Laws of Motion             $dp/dt = \mathbf{F}$

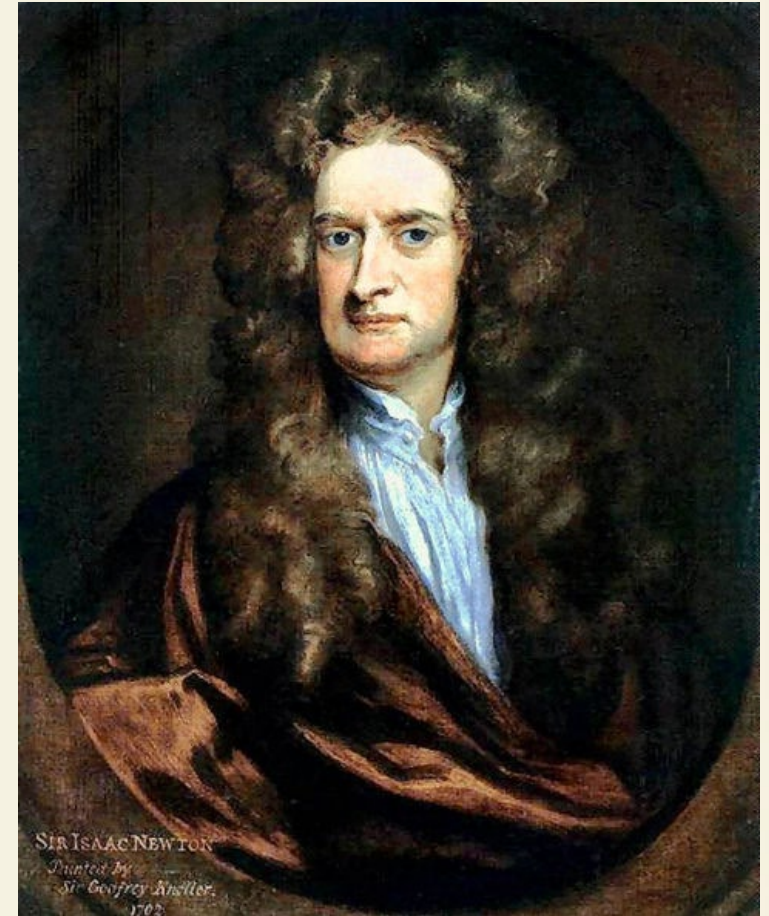
Law of Gravitation

Kepler's Laws Explained

Calculus (fluxions)

*... Space and time are absolute ...*

**Determinism** - Given  $\mathbf{F}$ , predict  $\mathbf{x}$  and  $\mathbf{v}$



## **Unification**

... the force responsible for bodies falling on the Earth is the same as that causing the moon to follow its orbit.

# Reformulations of $F = ma$

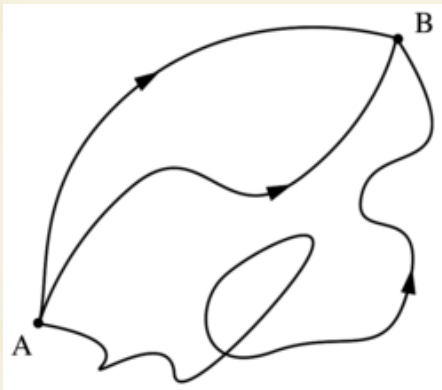
## From Classical Dynamics

Euler (1707-1783) **Variational Calculus**

D'Alembert (1717-1783) **Virtual Work**

Lagrange (1736-1813) **Lagrangian Mechanics**

Hamilton (1805-1865) **Hamiltonian Mechanics**



## Principles

Fermat's:	least time
d'Alembert's:	virtual work
Hamilton's:	least action

Define the action

$$S = \int_{t_1}^{t_2} L dt, \text{ for } L = T - V.$$

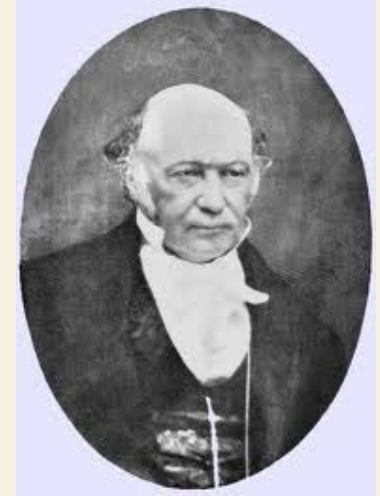
Require:  $\delta S = 0$ .

Then,  $[F = ma]$

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}} \right) - \frac{\partial L}{\partial x} = 0.$$

Optics	Vis viva	Action				Path Integrals
Fermat – Leibniz vs Maupertuis	Euler – Lagrange	Hamilton	.....	-	Feynman	
1662	1686	1744	1744	1788	1834	1948

# Hamilton's Formulation



Phase Space ( $q = x, p = m dx/dt$ ),

Initial  $(q, p)$  + 2<sup>nd</sup> Law  $\Rightarrow$  Motion for all  $t$

Ex: Free particle,  $p = \text{const}$   $E = \frac{p^2}{2m} + \frac{1}{2}kq^2$

Harmonic Oscillator,

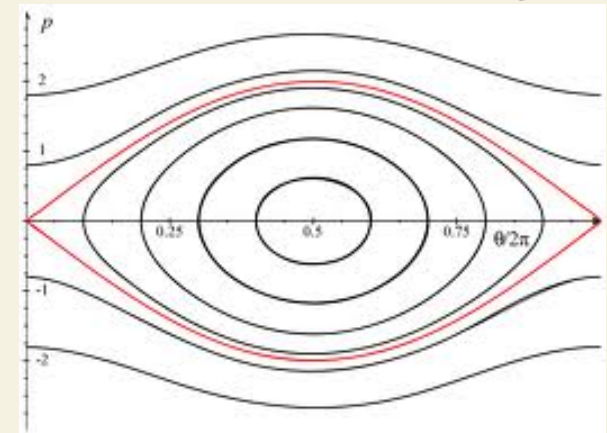
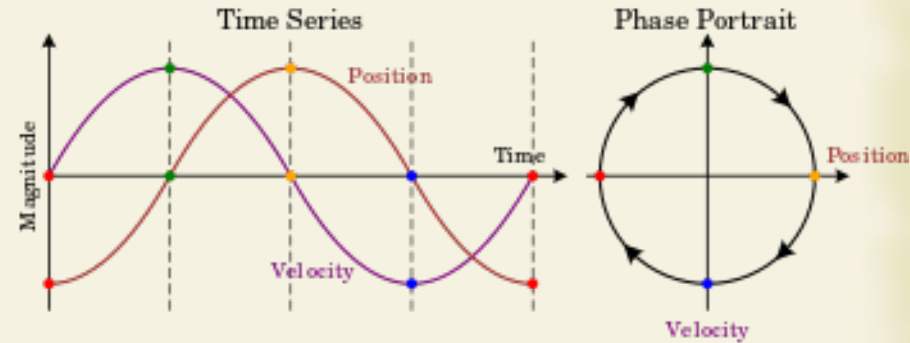
Energy Conservation  $E = \text{const}$

Hamiltonian  $H(p, q) = T(p) + V(q)$

Hamilton's Eqns  $\Leftrightarrow$  Newton's Laws

$$\frac{dq}{dt} = \frac{\partial H(q, p)}{\partial p}$$

$$\frac{dp}{dt} = -\frac{\partial H(q, p)}{\partial q}$$



Poisson: For  $F(p, q)$ ,  $\frac{d}{dt} F(p, q) = \frac{\partial F}{\partial q} \frac{\partial H}{\partial p} - \frac{\partial F}{\partial p} \frac{\partial H}{\partial q} = \{F, H\}$



# Electricity and Magnetism

Magnetism (Lode stones, Compasses)

Electricity (static, lightning)

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William Gilbert (1544-1603)

(amber, elektron, magnetic + electric forces different)

Thomas Browne (1605-1682) (“electricity”)

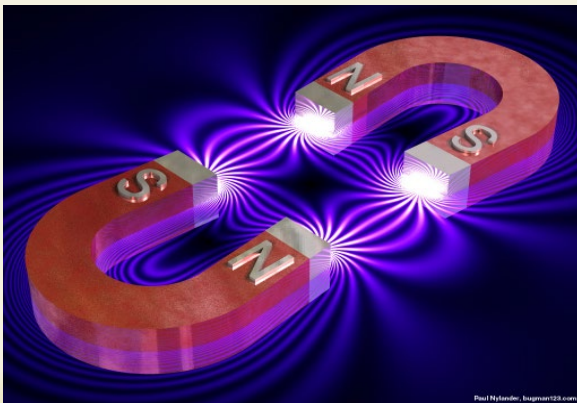
Benjamin Franklin (1706-1790) (“+/- electricity”)

Henry Cavendish (1731-1810)

Charles-Augustin de Coulomb (1736-1806)

Luigi Galvani (1737-1798) (animal electricity)

Alessandro Volta (1745-1827) (electrochemical cell)





# Electromagnetism

Hans Oersted (1777-1851)

1820 current deflects compass needles, made Al

André-Marie Ampère, (1775 - 1836)

“electrodynamics,” current carrying wires attract, telegraph

Georg Simon Ohm (1789-1854)

Ohm’s Law - 1827

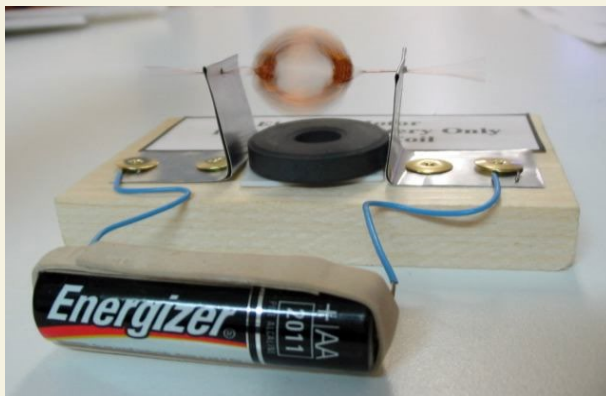
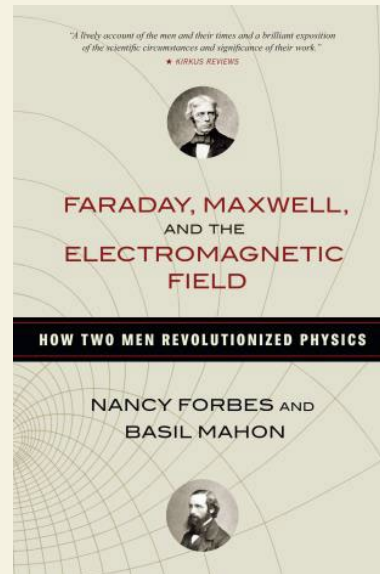
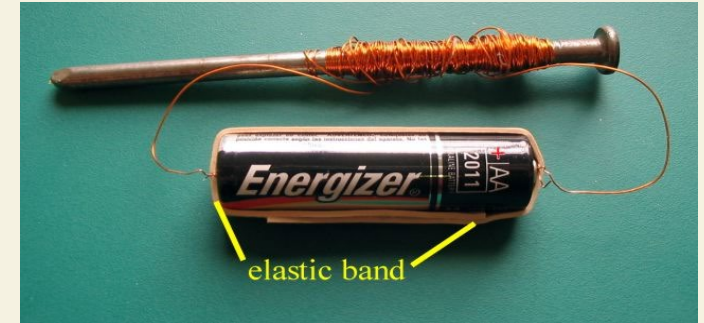
Joseph Henry (1797-1878)

electromagnetic induction, first motor, relays

Michael Faraday (1791-1867)

electrolysis, motors, induction coils, ...

Introduced concept of a field.



<http://hilaroad.com/camp/projects/magnet.html>

# Electromagnetic Waves

James Clerk Maxwell (1831-1879)

- Theory of electromagnetism – 1865.

1873, Maxwell also used the quaternions of Hamilton (1843),

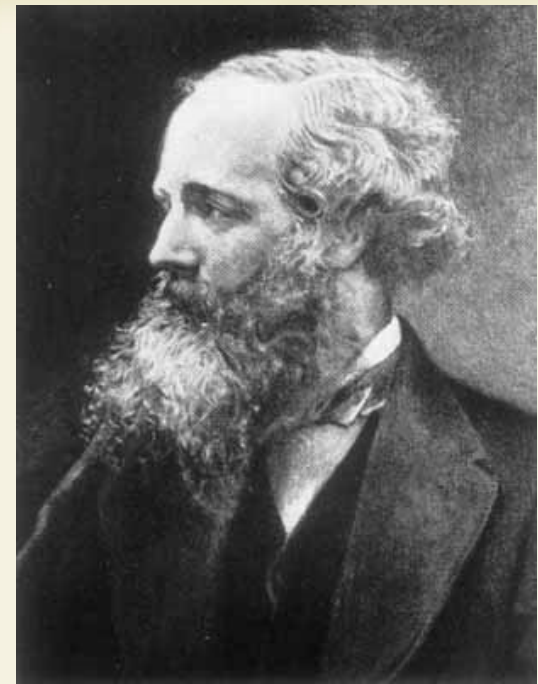
1880s, Heaviside reduced the 20 PDEs – 12 to 4,  
using symbolic vector calculus

[independent of Josiah Gibbs, *Vector Analysis*, (1881-1884)]

1890, Hertz presented other forms

- Predicted the electromagnetic waves - 1862.

Electromagnetic waves travel:  $c = 299,792,458 \text{ m/s} = 186,000 \text{ mi/s}$



## The Maxwellians

George Francis FitzGerald (1851–1901),

Oliver Lodge (1851–1940) and Oliver Heaviside (1850)–1925)

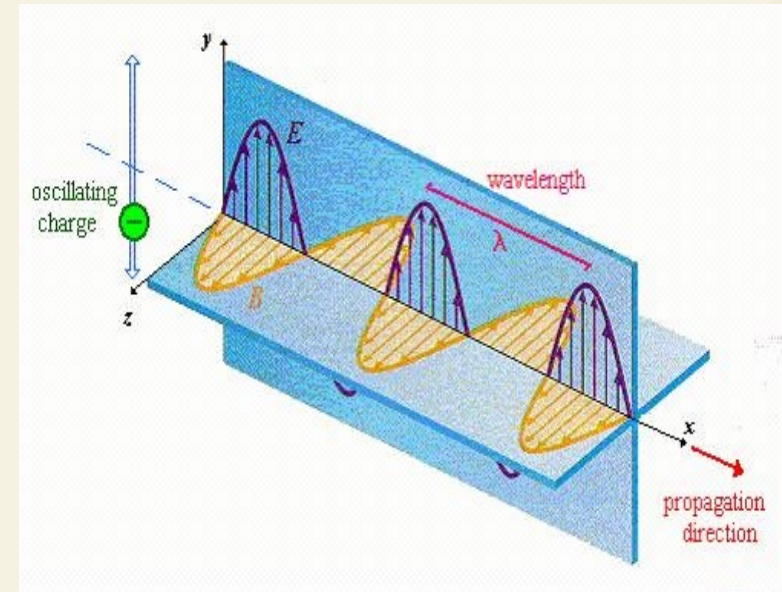
Heinrich Hertz (1857-1894)

- Sent the first radio waves – 1888.

- Marconi (1874-1937), practical radio waves – 1897

What is the medium? - *Luminiferous Aether*

(supported by Thomson, Stokes, ...)



# Maxwell's Equations

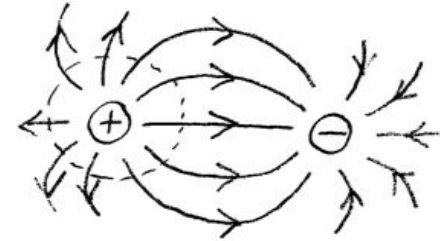
Not SI units!

## JAMES CLERK MAXWELL'S EQUATIONS

$$\nabla \cdot \vec{E} = 4\pi\rho$$

DIVERGENCE OF  $\vec{E}$       CHARGE DENSITY

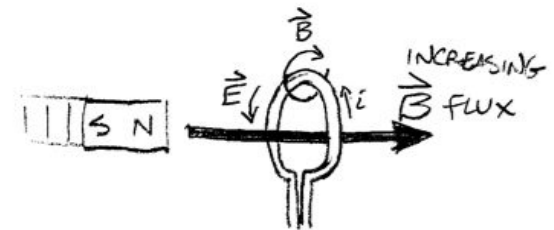
$\vec{E}$  DIVERGES OUT FROM POSITIVE CHARGES AND IN TOWARD NEGATIVE CHARGES. THE TOTAL FLUX OF  $\vec{E}$  THROUGH ANY CLOSED SURFACE IS PROPORTIONAL TO THE CHARGE INSIDE.



$$\nabla \times \vec{E} = -\frac{1}{c} \frac{d\vec{B}}{dt}$$

CURL OF  $\vec{E}$       SPEED OF LIGHT      RATE  $\vec{B}$  IS CHANGING

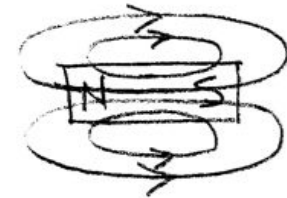
$\vec{E}$  CURLS AROUND CHANGING  $\vec{B}$  FIELDS (FARADAY'S LAW) IN A DIRECTION THAT WOULD MAKE A CURRENT THAT WOULD PRODUCE A  $\vec{B}$  FIELD TO OPPOSE THE CHANGE IN  $\vec{B}$  FLUX (LENZ'S LAW).



$$\nabla \cdot \vec{B} = 0$$

DIVERGENCE OF  $\vec{B}$

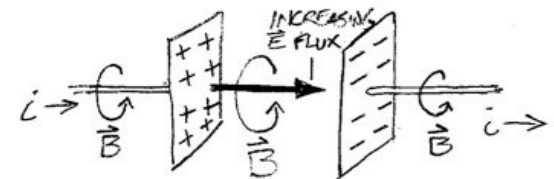
$\vec{B}$  NEVER DIVERGES. IT JUST LOOPS AROUND ON ITSELF.



$$\nabla \times \vec{B} = \frac{4\pi}{c} \vec{j} + \frac{1}{c} \frac{d\vec{E}}{dt}$$

CURL OF  $\vec{B}$       SPEED OF LIGHT      CURRENT DENSITY      RATE  $\vec{E}$  IS CHANGING

$\vec{B}$  CURLS AROUND CURRENTS AND CHANGES IN  $\vec{E}$  FIELDS





# Gaussian Units

## Unit Conversions



		Conversion	SI
Distance	cm	$10^{-2}$	m
Mass	g	$10^{-3}$	kg
Time	s	1	s
Force	dyne	$10^{-5}$	N
Energy	erg	$10^{-7}$	J
Power	erg/s	$10^{-7}$	W
Charge	esu	$3.336 \times 10^{-10}$	C
Electric Potential	statvolt	299.79	V
Magnetic Field	Gauss	$10^{-4}$	T

- $1 \text{ eV} = 1.6022 \times 10^{-12} \text{ erg} = 1.602 \times 10^{-19} \text{ J}$
- $1 \text{ Ry} = 13.6057 \text{ eV}$  (ionization energy of hydrogen)
- $1 \text{ C} = 2.9979 \times 10^9 \text{ esu}$ ,  $1 \text{ statcoul} = 1 \text{ esu}$
- $1 \text{ \AA} = 10^{-10} \text{ m}$
- $1 \text{ eV}/c^2 = 1.7827 \times 10^{-36} \text{ kg}$
- $(\mu_0 \epsilon_0)^{-1/2} = 299,792,458 \text{ m/s}$ ;  $(\mu_0/\epsilon_0)^{1/2} \approx 377 \Omega$

$$e^2 \text{ (Gaussian)} \rightarrow \frac{e^2}{4\pi\epsilon_0} \text{ (SI)}$$

$$F_{\text{Coulomb}} = \frac{q_1 q_2}{r^2}$$

# New Questions

- Waves – What is the medium?  
    Michelson-Morley (1887) - could not detect it.
- Spectroscopy – Why the spectral lines?
- Blackbody Spectrum – Describe the dependence on  $\lambda$ .
- Lorentz Invariance – Explain speed of light in moving media.

Led to Revolutions in Physics in the 1900s!

# WAVES



## What are waves?

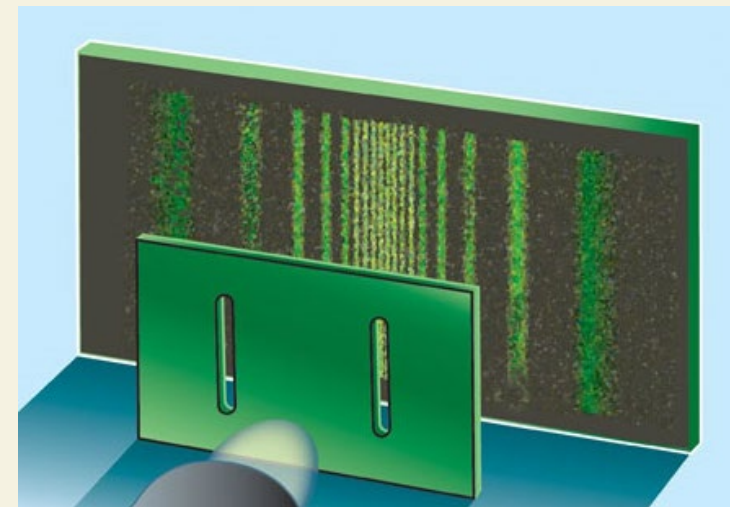
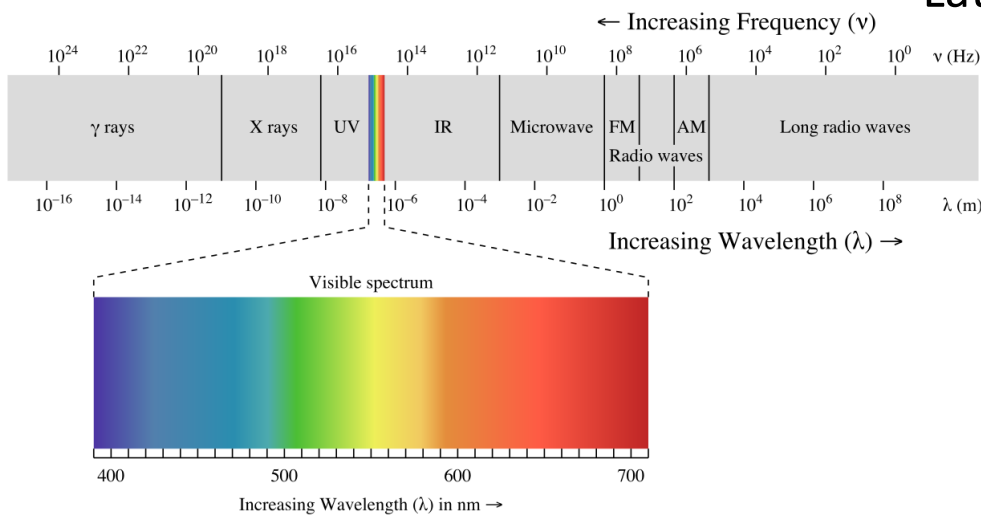
- **Characteristics**

- Wavelength, Frequency, Wavespeed

- **Behavior**

- Superposition, Interference. Diffraction
- Thomas Young, 1801, diffraction
- 1817, the Académie des Sciences: diffraction would be the topic for the biannual physics *Grand Prix*. – proposed by corpuscular theorists.
- Augustin-Jean Fresnel used Huygen's Principle, 1678.
- Later – Airy, Stokes, Helmholtz, Kirchoff , and others.

$$\frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = \nabla^2 E$$





# Spectroscopy

**Robert Bunsen and Gustav Kirchhoff**

developed the spectroscope, 1859.

Ionized gas gives off radiation

**Johann Balmer 1885**

Spectral Lines: Hydrogen  
410, 434, 486, 656 nm

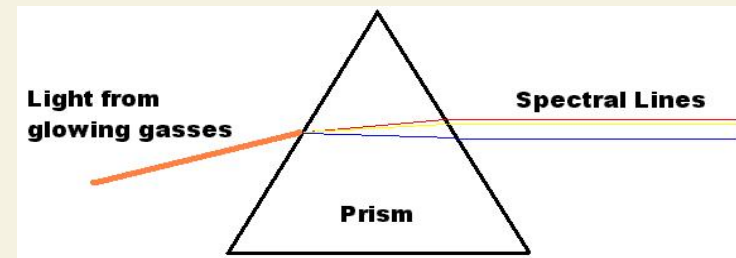
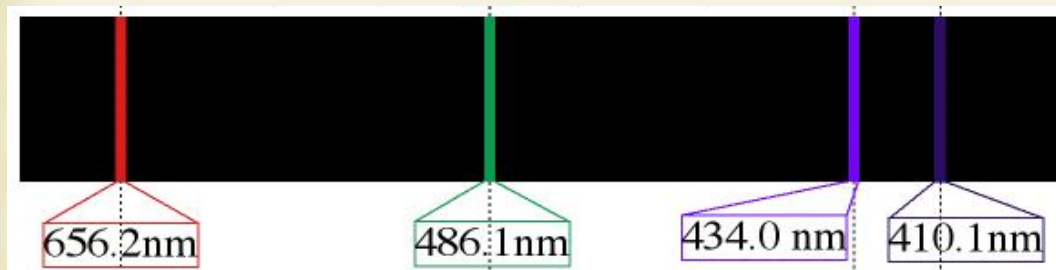
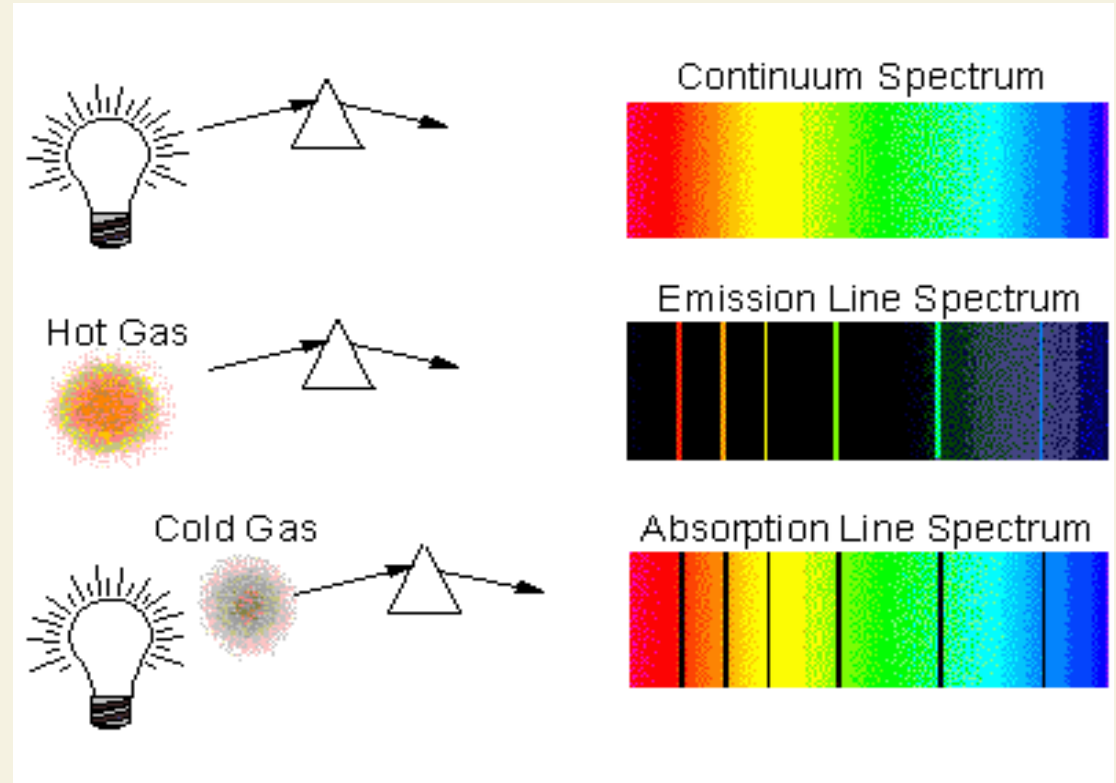
Derived Empirical Formula:

$$\lambda = R \left( \frac{1}{4} - \frac{1}{n^2} \right)$$

Predicted 5th-7th lines

**Lyman Series (1906-1914), ultraviolet**

**Paschen Series (1908), infrared**

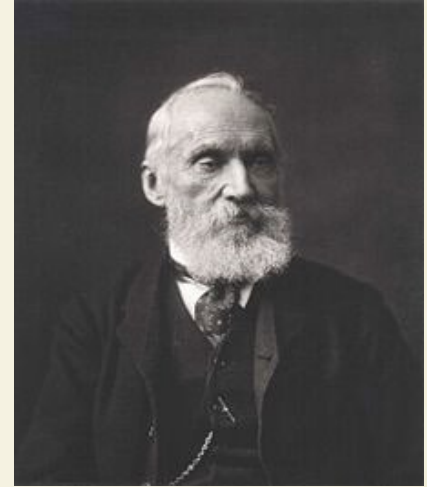


# Laws of Thermodynamics

**Engines: Watt, Carnot, Kelvin, Clausius**

**James Joule (1818-1889)**

Mechanical Equivalent of Heat



**Laws of Thermodynamics**

1. Adding heat energy or doing work on a body increases internal energy.  
(Energy conservation)
2. A body will not spontaneously get hotter.  
(Entropy and the Arrow of Time)

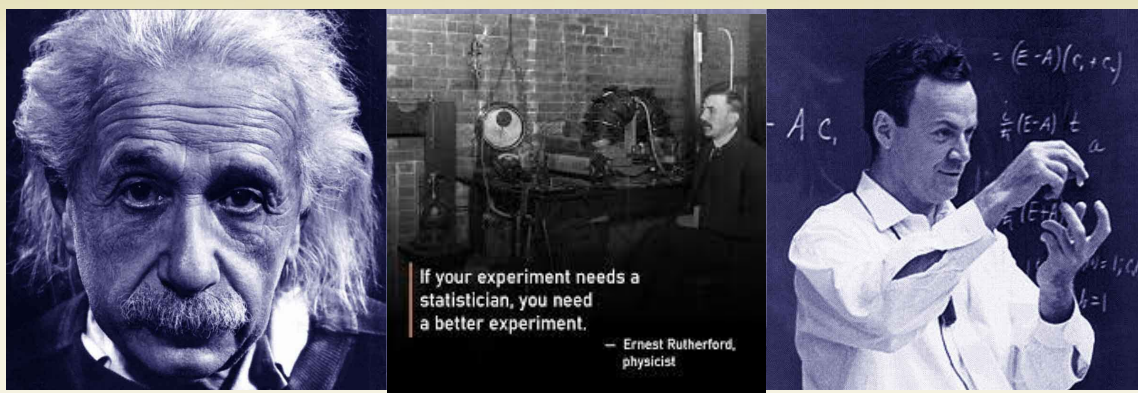
**Joseph Stefan (1835-1893) and Ludwig Boltzmann (1844-1906)**

**Heated bodies Radiate - Stefan-Boltzmann Law**

Radiation from blackbody proportional to  $T^4$ .

$$P = e\sigma AT^4$$

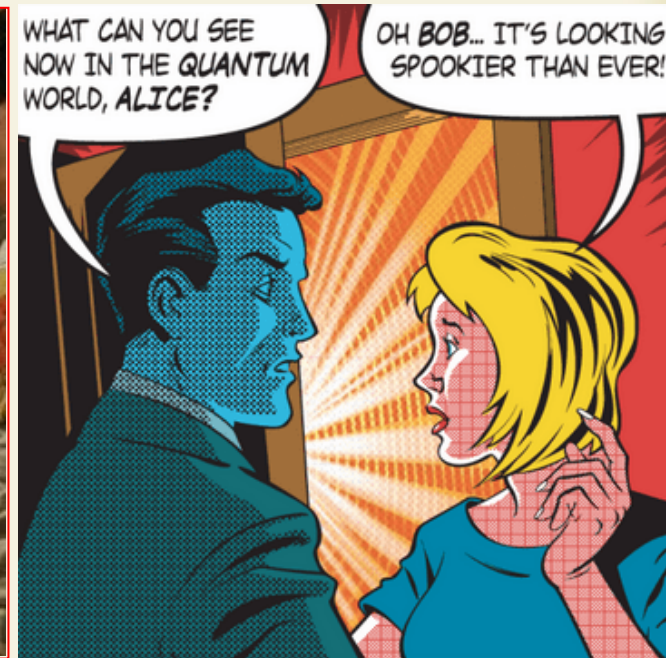
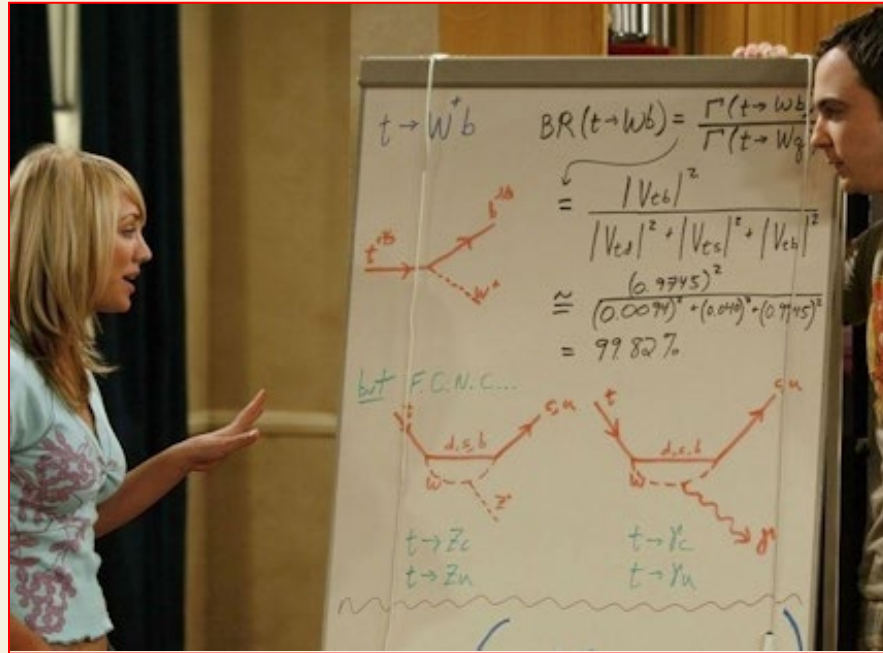
**Maxwell-Boltzmann Statistical Mechanics – *Bah Humbug!***



$$H|\psi\rangle = i\hbar \frac{\partial}{\partial t} |\psi\rangle$$

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

*This is not your grandfather's physics*



$$\left[ i\hbar \gamma^\mu \partial_\mu - mc \right] \psi = 0$$

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$



# Physics Revolutions

**Albert A. Michelson, (1852 – 1931)**

*“it seems probable that most of the grand underlying principles have been firmly established and that further advances are to be sought chiefly in the rigorous application of these principles .” - 1894*



1897 **Joseph (J.J.) Thomson** measures electron, NP 1906  
**“plum-pudding” model of atom**



# Radioactivity and the Atom

1895 - Wilhelm Röntgen discovers **X-rays**. NP 1901

1896 - Henri Becquerel discovers **radioactivity** NP 1903

1897 - J.J. Thomson discovers the **electron** NP 1906.

1898 - Marie and Pierre Curie discover the first radioactive elements:  
**radium and polonium** NP 1903.

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1899 - Ernest Rutherford divided radiation into **alpha and beta rays** NP 1908.

1900 - Pierre Curie observes **gamma rays**.

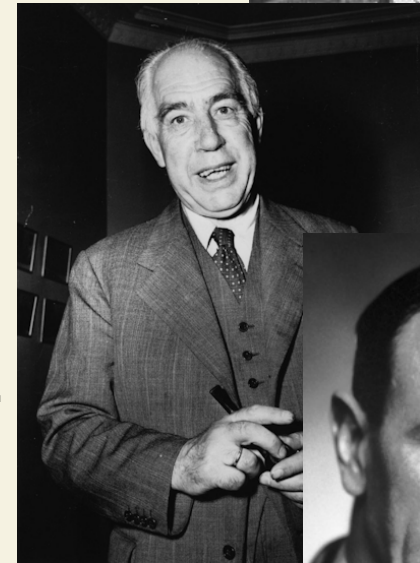
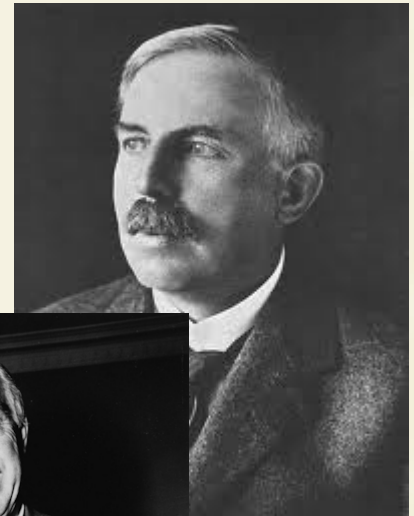
1911 - Ernest Rutherford discovers the **atomic nucleus**

1913 - Niels Bohr introduces the **first atomic model**, NP 1922  
the mini solar system.

1913 - Hans Geiger invents **counter** for measuring radioactivity.

1920 - Ernest Rutherford discovered and named the **proton**.

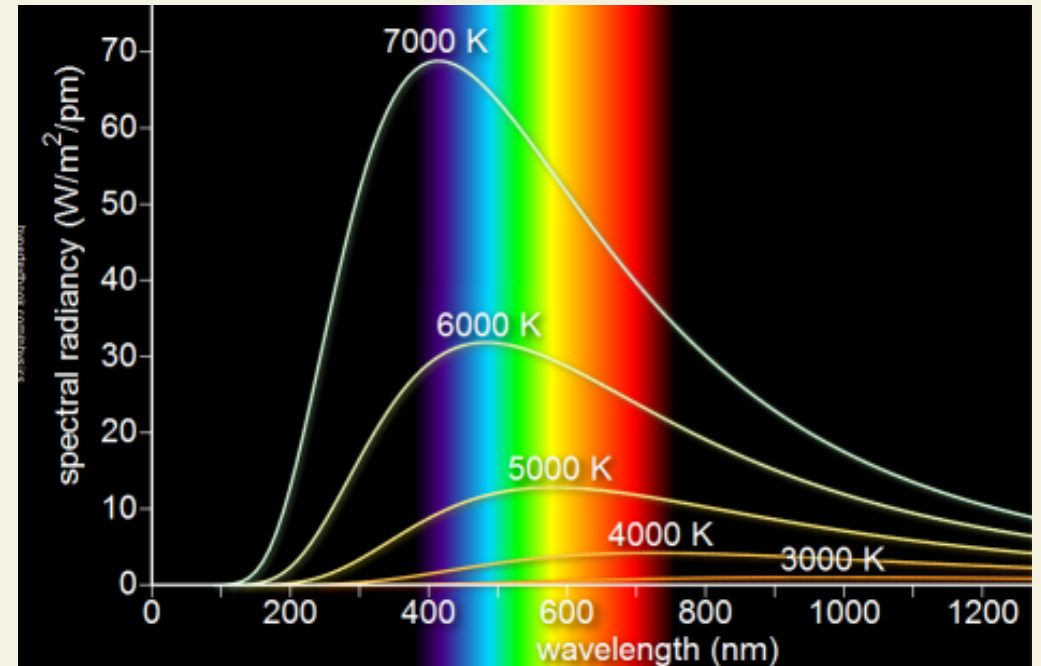
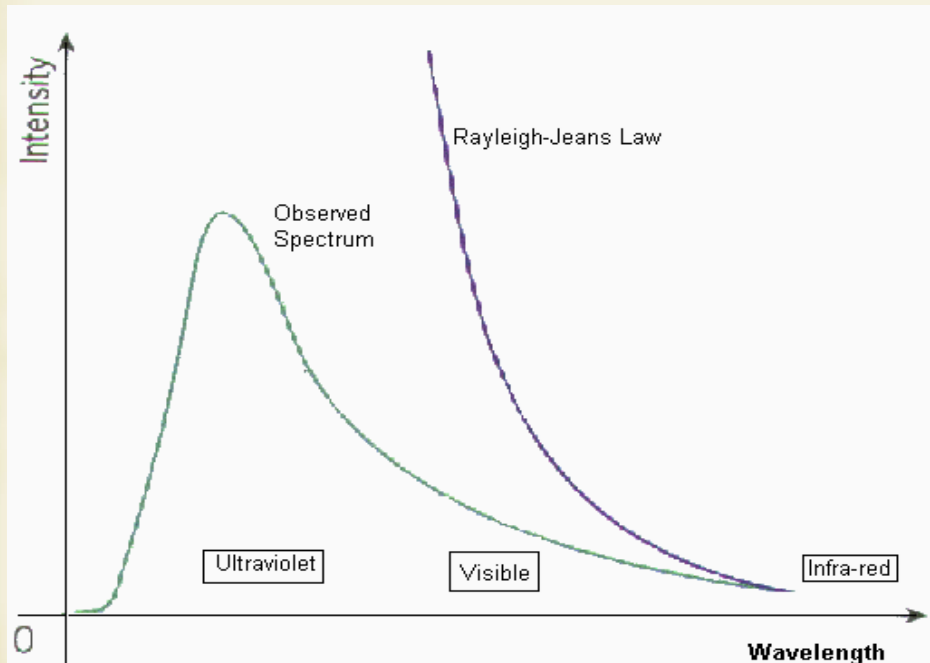
1932 - James Chadwick discovers the **neutron**. NP 1935



# Blackbody Spectrum

**Blackbody** - a theoretical object that absorbs 100% of the radiation that hits it.

**Wien's Law** (1896) <sup>NP 1911</sup> and **Rayleigh** <sup>NP 1904</sup> - **Jeans Law** (1900/1905)



**Ultraviolet (Rayleigh–Jeans) Catastrophe** “... when you turn on your toaster, you are instantly fried by a massive gamma ray burst, since your little blackbody toaster should emit infinite energy at the shortest wavelengths.”

*For fixed  $T$ , monochromatic energy density becomes infinite at infinitely small wavelengths!*



# Quantum Theory

**Max Planck** NP 1918

(Karl Ernst Ludwig Max Planck 1858-1947)

What makes hot solids glow different colors?

Nov. 1900 – he used Boltzmann's theory.

*oscillators can only vibrate at discrete frequencies:*

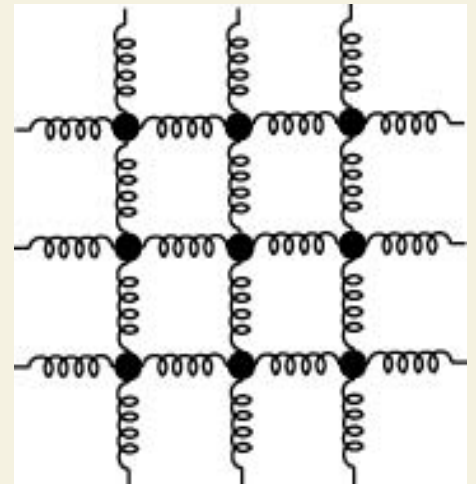
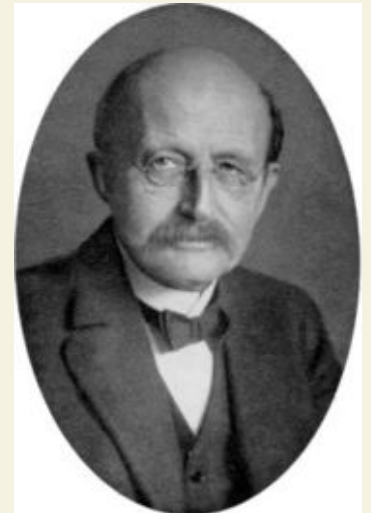
$$E_n = n(hf), n = 1, 2, 3 \dots$$

Thus, the energy difference

$$\Delta E = hf,$$

where Planck's constant is given by

$$h = 6.63 \times 10^{-34} \text{ Js}$$



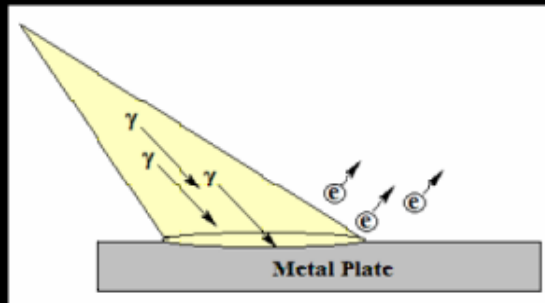
# Albert Einstein - 1905

Received 18 March and published 9 June,

## Photoelectric Effect

Light can cause currents

- **Electrons can be ejected from irradiated metal plates.**
- **Light can be act like either particles (quanta) or waves.**
- **Extended Planck's ideas of energy quantization.**
- **Lead to explanation of electromagnetic spectra,**
- **Lead to the development of lasers, transistors and other applications.**

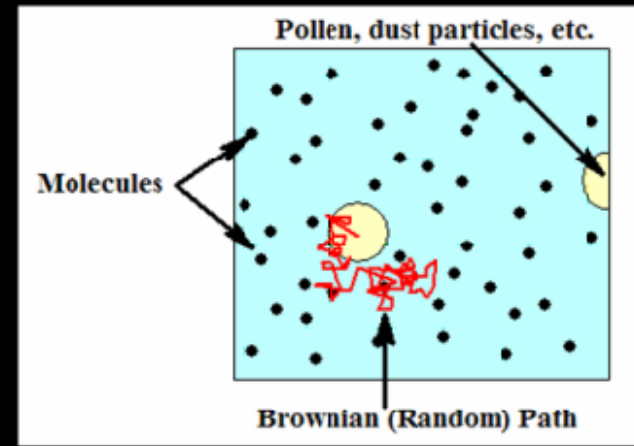


Received 11 May and published 18 July

## Brownian Motion

*the random movement of particles suspended in a fluid*

- **Explained the observations credited to Robert Brown, 1827**
- **Predicted molecular motion and size through the effects of collisions with larger particles**
- **Einstein's work lead to an acceptance of molecular theory**



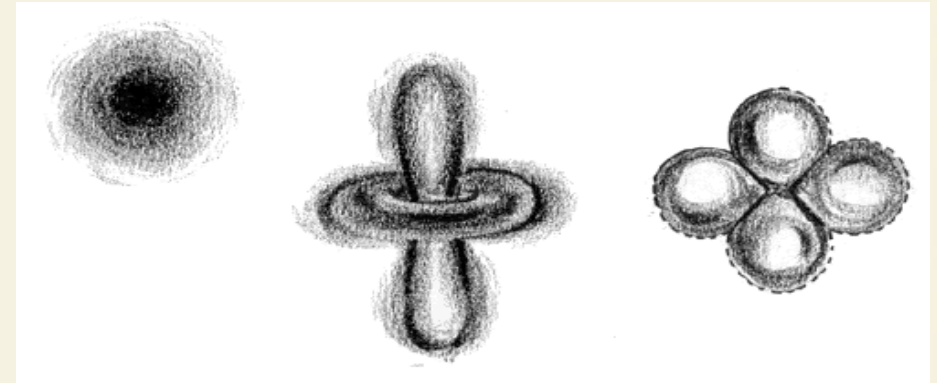
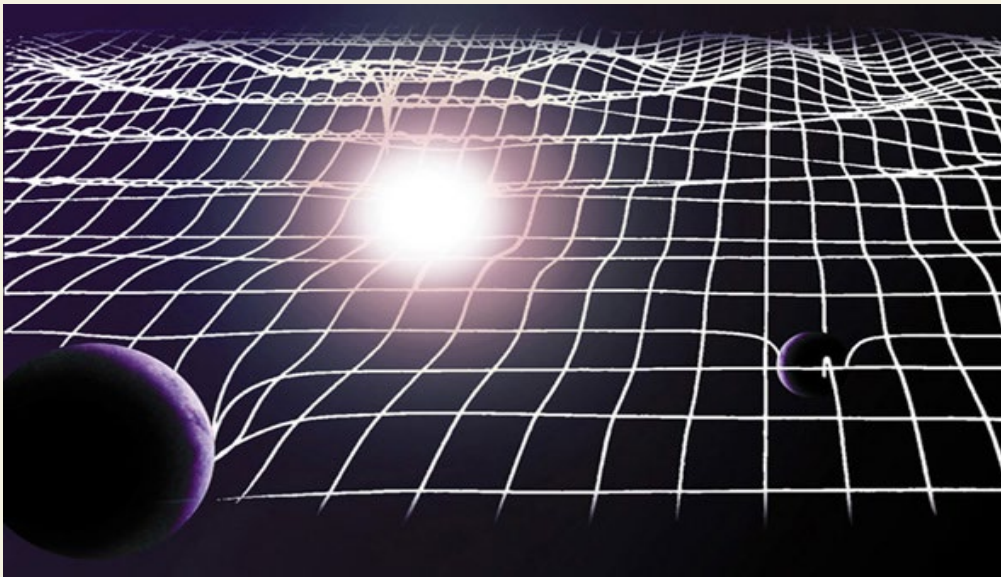
Special Relativity: received on 30 June and published 26 September

Mass-Energy Equivalence: received September 27 and published 21 November

# Paradigm Shifts – *in progress*

## Relativity

Space and Time are not absolute, there is no preferred frame.  
and not Euclidean



## Quantum Mechanics

Loss of Determinism:

Determinism: If we knew all of the initial conditions, we couldn't predict the exact position and velocity of an electron.



$$\Delta t = \gamma \Delta \tau$$

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$



# Bohr's Atom - 1913

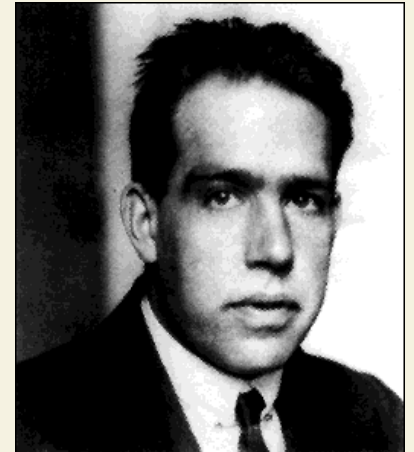
**Niels Bohr (1885-1962)**

Accelerating electrons radiate at specific energies.

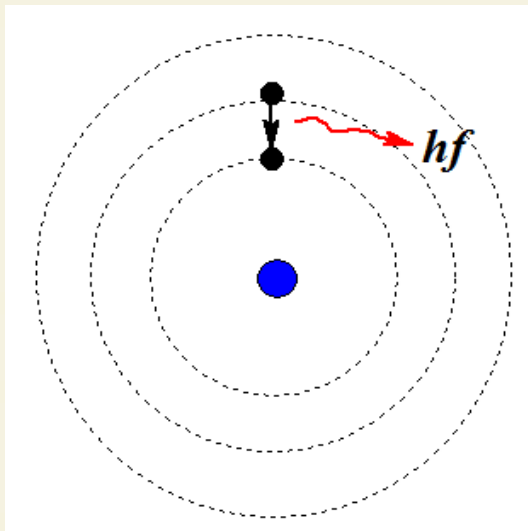
Are atoms stable?

Assume angular momentum is quantized.

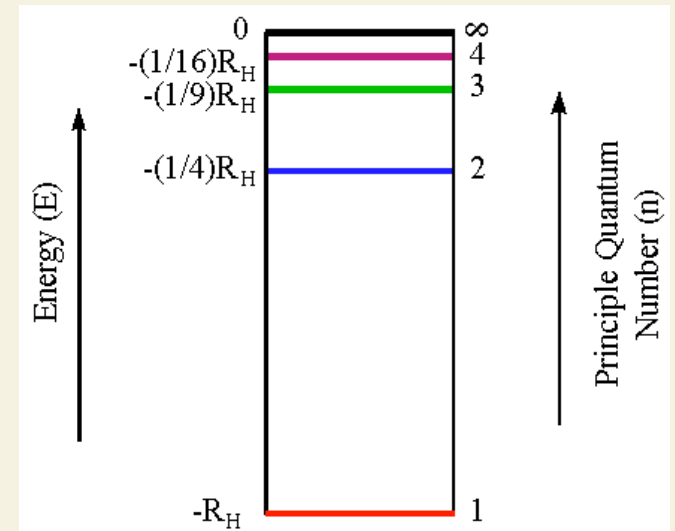
Derived Balmer's formula.



Niels Bohr



$$E_n = R_H \left( \frac{1}{m^2} - \frac{1}{n^2} \right)$$



# Early History - Quantum Mechanics

1900 - Planck Explains Blackbody Radiation NP 1906

1905 - Einstein - the Photoelectric Effect, Photons NP 1921

1913 - The Bohr Model for Hydrogen NP 1922

1916 - Confirmation of photon, Millikan NP 1923

1922 - Stern-Gerlach Experiment NP 1943

1923 - Compton NP 1927 Effect - X-Ray Scattering

1924 - de Broglie NP 1929 - Particles Behave Like Waves

1925 - Matrix Mechanics - Heisenberg NP 1932, Born NP 1954, Jordan  
- Pauli Principle NP 1935

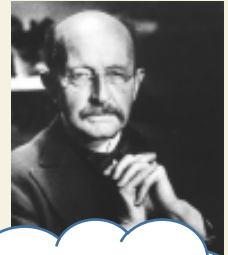
- Uhlenbeck and Goudsmit, spinning particles

1926 - Wave Mechanics - Schrödinger NP 1933

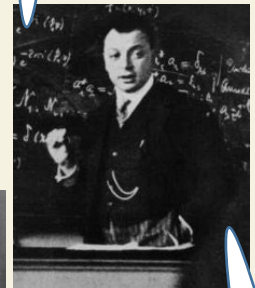
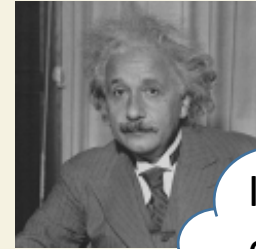
1927 - The Uncertainty Principle - Heisenberg

- Davisson NP 1937 - Germer, Thomson NP 1937 - Verified de Broglie's idea

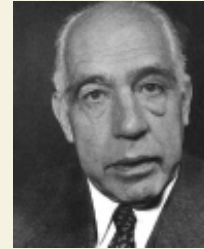
1928 - Relativistic Quantum Mechanics - Dirac NP 1933



I introduced doubling e-states in 1924



I gave theory of spin in 1927



$i\gamma\partial\psi = m\psi$

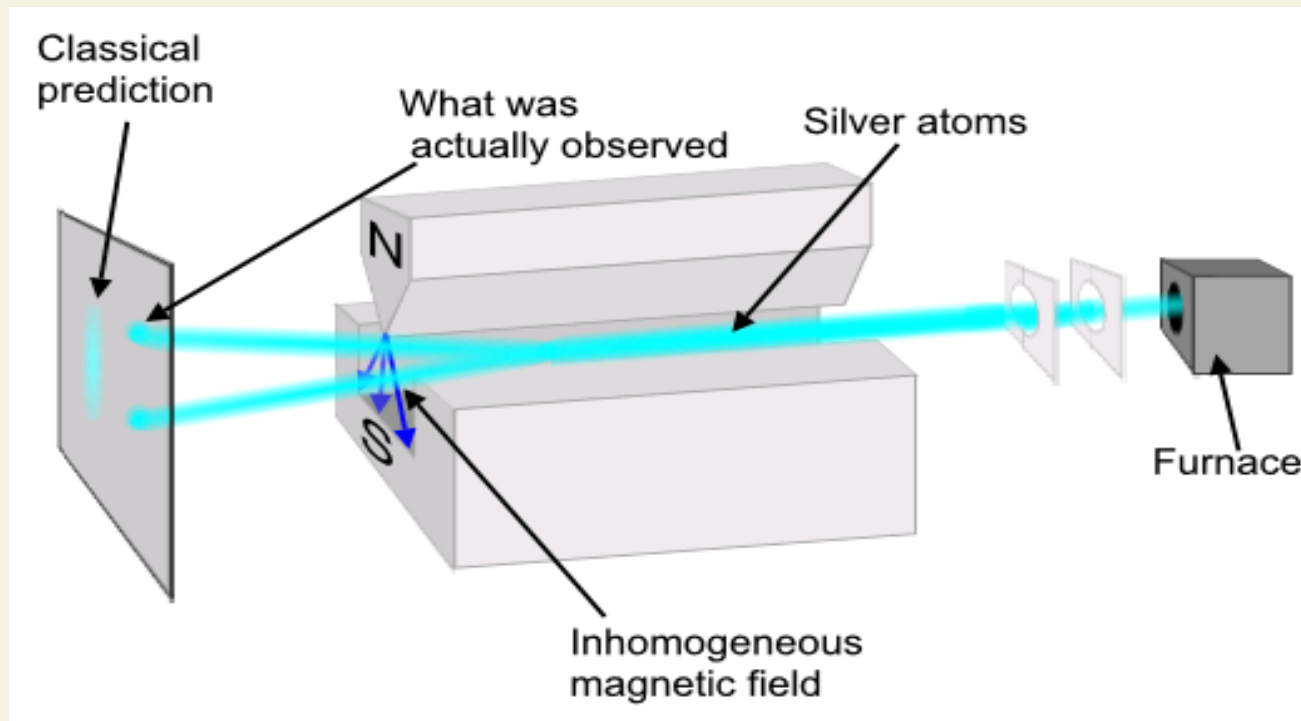
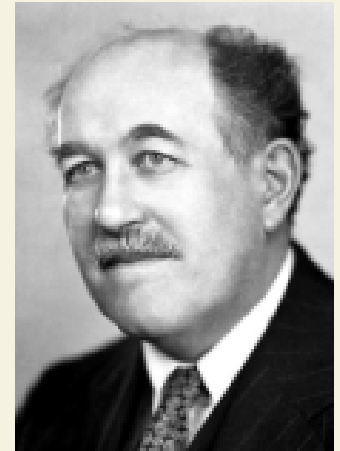
# Stern-Gerlach Experiment - 1922

Otto Stern (1888-1969)

Walther Gerlach (1889-1979) – Was it the cigar?

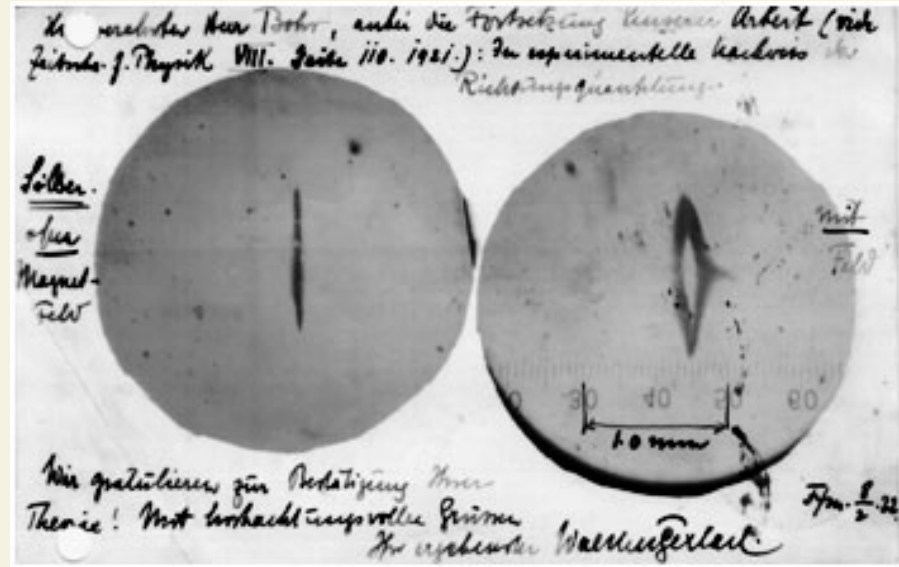
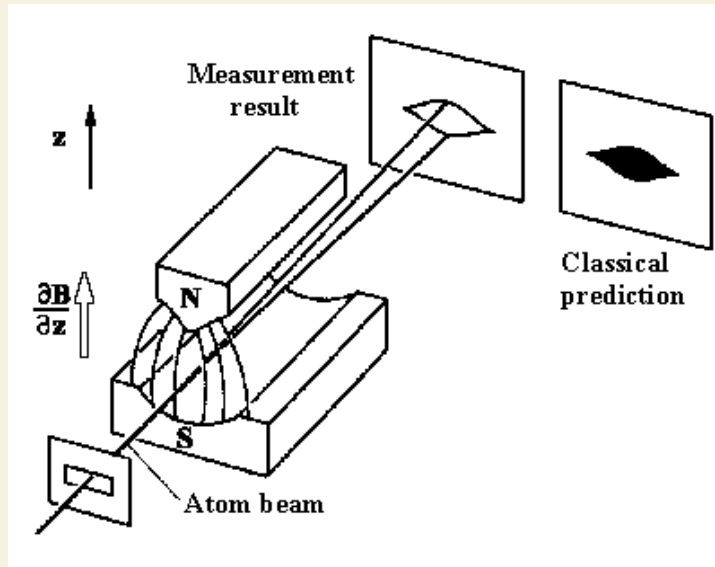
Ag atoms sent through inhomogeneous magnetic field.

Demonstration of space quantization.





# Stern-Gerlach Results



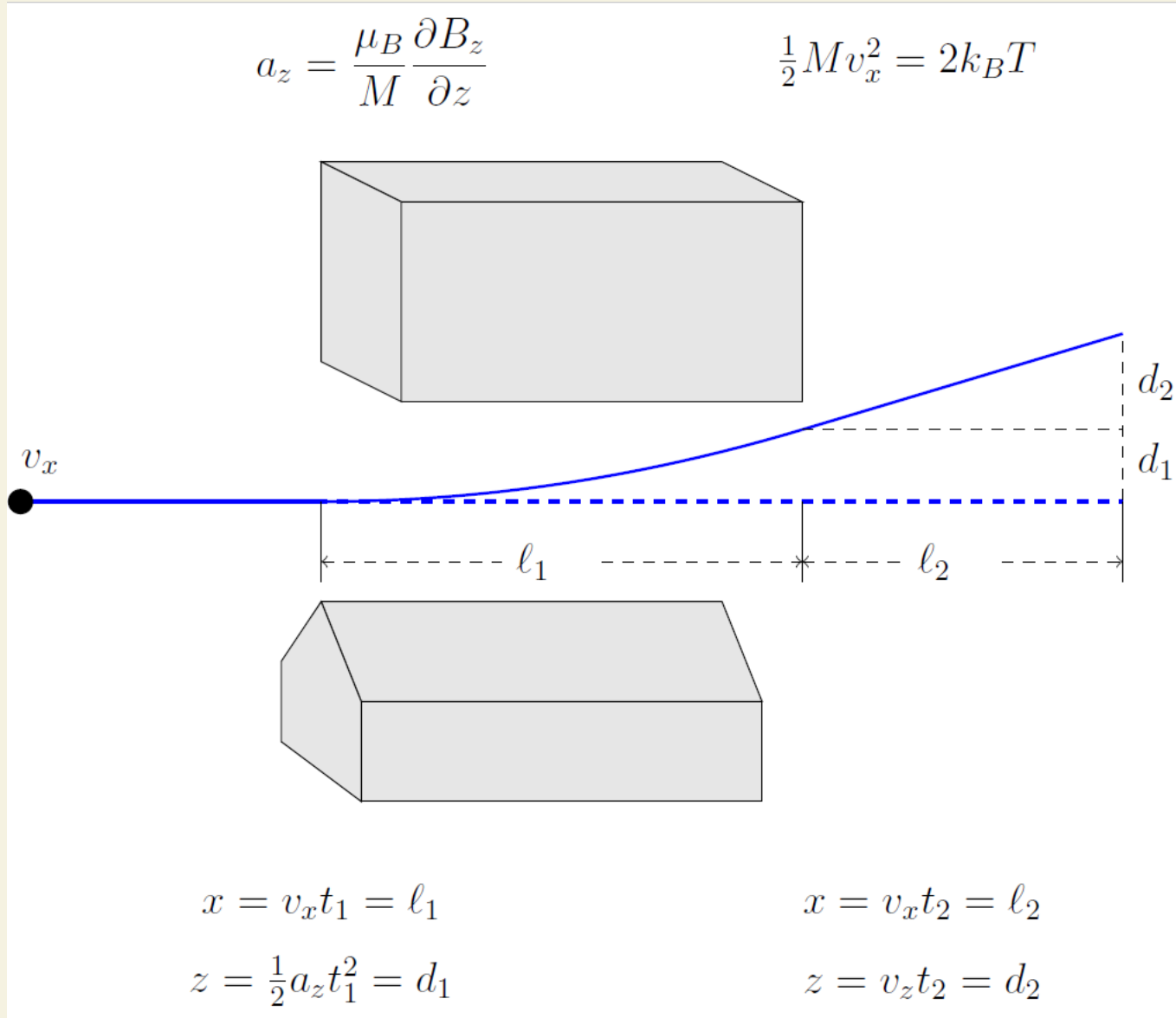
Test of classical vs quantum theory of angular momentum ( $L$ ).

$L = 0$  or  $L = 1$  – No splitting for  $L = 0$ ?

Uhlenbeck and Goldsmit (1925,1926) proposed **intrinsic spin**.

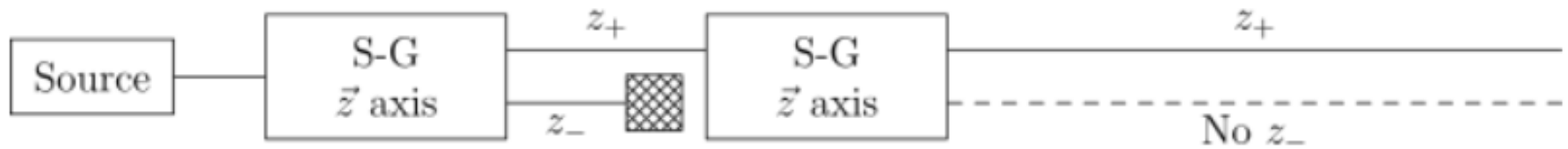
- to explain the anomalous Zeeman effect,  
(the splitting of spectral lines in a magnetic field).

# Stern-Gerlach Particle Path – Problem 1

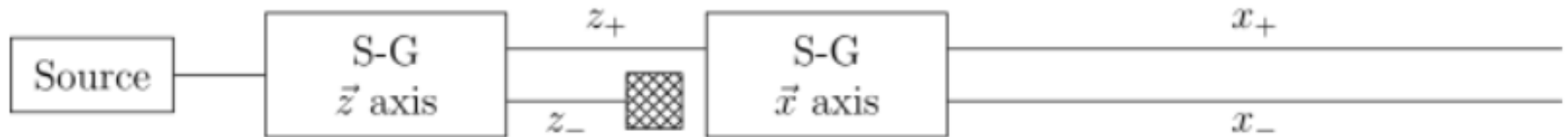


# Stern-Gerlach Thought Experiments

1



2



3

